PHYS 3446 – Lecture #9

Monday, Oct. 3, 2016 Dr. **Jae** Yu

- Nuclear Models
 - Fermi Gas Model
 - Shell Model
 - Collective Model
 - Super-deformed Nuclei
- Nuclear Radiation
 - Alpha decay
 - Beta decay



Announcement

- First term exam
 - Date and time: 2:30 3:50pm, Monday, Oct. 10
 - Location: SH125
 - Covers: Ch 1 Ch 3 or what we finish Wednesday, Oct. 5, + Appendix A
 - Can bring your calculator but no phone or computer can be used as a replacement
- First quiz results
 - Class average: 56.5/120
 - Equivalent to 47.1/100
 - Class top score: 89/120



- An early attempt to incorporate quantum effects
- Assumes nucleus as a gas of free protons and neutrons confined to the nuclear volume
 - The nucleons occupy quantized (discrete) energy levels
 - Nucleons are moving inside a spherically symmetric well with the range determined by the radius of the nucleus
 - Depth of the well is adjusted to obtain correct binding energy
- Protons carry electric charge → Senses slightly different potential than neutrons



- Nucleons are Fermions (spin 1/2 particles) so
 - Obey Pauli exclusion principle
 - Any given energy level can be occupied by at most two identical nucleons – opposite spin projections
- For a greater stability, the energy levels fill up from the bottom up to the Fermi level
 - Fermi level: Highest, fully occupied energy level (E_F)
- Binding energies are given as follows:
 - BE of the last nucleon= E_F since no Fermions above E_F
 - In other words, the level occupied by Fermion reflects the BE of the last nucleon



- Experimental observations show BE is charge independent
- If the well depth is the same for p and n, BE for the last nucleon would be charge dependent for heavy nuclei (Why?)
 - Since there are more neutrons than protons, neutrons sit higher E_F



Same Depth Potential Wells

Neutron Well

Proton Well





- Experimental observations show BE is charge independent
- If the well depth is the same for p and n, BE for the last nucleon would be charge dependent for heavy nuclei (Why?)
 - Since there are more neutrons than protons, neutrons sit higher E_F
 - But experiments observed otherwise
- $E_{\rm F}$ must be the same for protons and neutrons. How do we make this happen?
 - Make protons move to a shallower potential well
- What happens if this weren't the case?
 - Nucleus is unstable.
 - All neutrons at higher energy levels would undergo a β -decay and transition to lower proton levels



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Fermi Gas Model: E_F vs n_F

- Fermi momentum: $E_F = p_F^2 / 2m \, r \, p_F = \sqrt{2mE_F}$
- Volume for momentum space up to Fermi level $V_{p_F} = \frac{4\pi}{2} p_F^3$
- Total volume for the states (kinematic phase space)
 - Proportional to the total number of quantum states in the system

$$V_{TOT} = V \cdot V_{p_F} = \frac{4\pi}{3} r_0^3 A \cdot \frac{4\pi}{3} p_F^3 = \left(\frac{4\pi}{3}\right)^2 A \left(r_0 p_F\right)^3$$

- Using Heisenberg's uncertainty principle: $\Delta x \Delta p \ge \frac{1}{2}\hbar$
- The minimum volume associated with a physical system becomes $V_{state} = (2\pi\hbar)^3$
- The n_F that can fill up to E_F is

$$n_{F} = \frac{V_{TOT}}{(2\pi\hbar)^{3}} = \frac{2}{(2\pi\hbar)^{3}} \left(\frac{4\pi}{3}\right)^{2} A(r_{0}p_{F})^{3} = \frac{4}{9\pi} A\left(\frac{r_{0}p_{F}}{\hbar}\right)^{3}$$

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Fermi Gas Model: E_F vs n_F

 Let's consider a nucleus with N=Z=A/2 and assume that all states up to Fermi level are filled

$$N = Z = \frac{A}{2} = \frac{4}{9\pi} A \left(\frac{r_0 p_F}{\hbar}\right)^3 \qquad \text{or} \qquad p_F = \frac{\hbar}{r_0} \left(\frac{9\pi}{8}\right)^{1/3}$$

- What do you see about p_F above?
 - Fermi momentum is constant, independent of the number of nucleons

$$E_F = \frac{p_F^2}{2m} = \frac{1}{2m} \left(\frac{\hbar}{r_0}\right)^2 \left(\frac{9\pi}{8}\right)^{2/3} \approx \frac{2.32}{2mc^2} \left(\frac{\hbar c}{r_0}\right)^2 \approx \frac{2.32}{2 \cdot 940} \left(\frac{197MeV - fm}{1.2fm}\right) \approx 33MeV$$

- Using the average BE of -8MeV, the depth of potential well (V_0) is ~40MeV
 - Consistent with other findings
- This model is a natural way of accounting for a₄ term in Bethe-Weizsacker mass formula



Nuclear Models: Shell Model

- Exploit the success of atomic model
 - Uses orbital structure of nucleons
 - Electron energy levels are quantized
 - Limited number of electrons in each level based on the available spin and angular momentum configurations
 - For nth energy level, *l* angular momentum (*l*<n), one expects a total of 2(2*l*+1) possible degenerate states for electrons



Atomic Shell Model Reminder

- Orbits and energy levels an electron can occupy are labeled by
 - Principle quantum number: n
 - *n* can only be integer
 - For given n, energy degenerate orbital angular momentum: ℓ
 - The values are given from 0 to n-1 for each n
 - For any given orbital angular momentum, there are (2*l*+1) sub-states: m_l
 - $m_{\ell} = -\ell, -\ell+1, \ldots, 0, 1, \ldots, \ell-1, \ell$
 - Due to rotational symmetry of the Coulomb potential, all these sub-states are degenerate in energy
 - Since electrons are fermions w/ intrinsic spin angular momentum $\frac{1}{2}\hbar$,
 - Each of the sub-states can be occupied by two electrons
 - So the total number of state is $2(2\ell+1)$



Nuclear Models: Shell Model

- Exploit the success of atomic model
 - Uses orbital structure of nucleons
 - Electron energy levels are quantized
 - Limited number of electrons in each level based on available spin and angular momentum configurations
 - For nth energy level, *l* angular momentum (*l*<n), one expects a total of 2(2*l*+1) possible degenerate states for electrons
- Quantum numbers of individual nucleons are taken into account to affect the fine structure of spectra



Nuclear Models: Shell Model

- Nuclei have magic numbers just like inert atoms
 - Atoms: Z=2, 10, 18, 36, 54
 - Nuclei: N=2, 8, 20, 28, 50, 82, and 126; Z=2, 8, 20, 28, 50, and 82
 - Magic Nuclei: Nuclei with either N or Z a magic number
 Stable
 - Doubly magic nuclei: Nuclei with both N and Z magic numbers → Particularly stable (some examples?)
- Explains well the stability of nucleus



Shell Model: Various Potential Shapes

 To solve the equation of motion in quantum mechanics, Schrödinger equation, one must know the shape of the potential

$$- \left(\vec{\nabla}^2 + \frac{2m}{\hbar^2} \left(E - V(r)\right)\right) \varphi(\vec{r}) = 0$$

- Details of nuclear potential not well known
- A few shapes of potential energies tried out
 - Infinite square well: Each shell can contain up to 2(2*l*+1) nucleons



Nuclear Models: Shell Model – Square well potential case

N _M	n	l=n-1	N _s =2(2 <i>l</i> +1)	N _T
2	1	0	2	2
8	2	0,1	2+6	8
20	3	0,1,2	2+6+10	18
28	4	0,1,2,3	2+6+10+14	32
50	5	0,1,2,3,4	2+6+10+14+18	50
82	6	0,1,2,3,4,5	2+6+10+14+18+22	72



Shell Model: Various Potential Shapes

- To solve equation of motion in quantum mechanics, Schrödinger equation, one must know the shape of the potential
 - $-\left(\vec{\nabla}^2 + \frac{2m}{\hbar^2} (E V(r))\right) \varphi(\vec{r}) = 0$
 - Details of nuclear potential not well known
- A few models of potential tried out
 - Infinite square well: Each shell can contain up to 2(2*l*+1) nucleons
 - Can predict 2, 8 and 50 but no other magic numbers
 - Three dimensional harmonic oscillator: $V(r) = \frac{1}{2}m\varpi^2 r^2$
 - Predicts 2, 8, 20, 40 and 70 → Some magic numbers predicted



Shell Model: Spin-Orbit Potential

- Central potential could not reproduce all magic numbers
- In 1940, Mayer and Jesen proposed a central potential + strong spin-orbit interaction w/

$$V_{TOT} = V(r) - f(r)\vec{L}\cdot\vec{S}$$

- $f(\mathbf{r})$ is an arbitrary empirical function of radial coordinates and chosen to fit the data
- For the spin-orbit interaction with the properly chosen f(r), a finite square well can split
- Reproduces all the desired magic numbers

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Predictions of the Shell Model

- Spin-Parity of large number of odd-A nuclei predicted well
 - Nucleons are Fermions and obey Pauli exclusion principle
 - − → Fill up ground state energy levels in pairs
 - Ground state of all even-even nuclei have zero total angular momentum
- The shell model cannot predict stable odd-odd nuclei spins
 - No prescription for how to combine the unpaired proton and neutron spins



Predictions of the Shell Model

• Magnetic Moment of neutron and proton are

$$\mu_p \approx 2.79 \mu_N \qquad \mu_n \approx -1.91 \mu_N$$

- Intrinsic magnetic moment of unpaired nucleons contribute to total magnetic moment of nuclei
 - What does a deuteron consist of?

$$\mu_D = \mu_p + \mu_n = 2.79 \,\mu_N - 1.91 \,\mu_N = 0.88 \,\mu_N$$

- Measured value is $\mu_D = 0.86 \,\mu_N$
- For Boron (¹⁰B⁵), the 5 neutrons and 5 protons have the same level structure: $(1S_{1/2})^2(1P_{3/2})^3$, leaving one of each unpaired proton and neutron in angular momentum *l*=1 state $\Rightarrow \mu = \frac{e\hbar}{2m_Nc} \cdot l = \frac{e\hbar}{2m_Nc} \cdot 1 = \mu_N$ $\mu_B = \mu_p + \mu_n + \mu_{orbit} = 2.79 \,\mu_N - 1.91 \mu_N + \mu_N = 1.88 \,\mu_N$
 - Measured value is $\mu_B = 1.80 \,\mu_N$
- Does not work well with heavy nuclei

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Collective Model

- For heavy nuclei, shell model predictions do not agree with experimental measurements
 - Especially in magnetic dipole moments
- Measured values of quadrupole moments for closed shells differ significantly with experiments
 - Some nuclei's large quadrupole moments suggests significant nonspherical shapes
 - The assumption of rotational symmetry in shell model does not seem quite right
- These deficiencies are somewhat covered through the reconciliation of liquid drop model and the Shell model
 - Bohr, Mottelson and Rainwater's collective model, 1953



Collective Model

- Assumption
 - Nucleus consists of hard core of nucleons in filled shells
 - Outer valence nucleons behave like the surface molecules in a liquid drop
 - Non-sphericity of the central core is caused by the surface motion of the valence nucleons
- Thus, in the collective model, the potential is a shell model with a spherically asymmetric potential
 - Aspherical nuclei can produce additional energy levels upon rotation while spherical ones cannot
- Important predictions of the collective model:
 - Existence of rotational and vibrational energy levels in nuclei
 - Accommodate decrease of spacing between first excite state and the ground level for even-even nuclei as A increases, since moment of inertia increases with A
 - Spacing is largest for closed shell nuclei, since they tend to be spherical

Super-deformed Nuclei

- Nuclei tend to have relatively small intrinsic spins
- Particularly stable nuclei predicted for A between 150 and 190
 with spheroidal character
 - Semi-major axis about a factor of 2 larger than semi-minor
- Heavy ion collisions in late 1980s produced super-deformed nuclei with angular momentum of $60\hbar$
- The energy level spacing of these observed through photon radiation seem to be fixed
- Different nuclei seem to have identical emissions as they spin down
- Problem with collective model and understanding of strong pairing of nucleon binding energy
- Understanding nuclear structure still in progress

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- Represents the disintegration of a parent nucleus to a daughter through an emission of a He nucleus
- Reaction equation is

$${}^{A}X^{Z} \rightarrow {}^{A-4}Y^{Z-2} + {}^{4}He^{2}$$

- α -decay is a spontaneous fission of the parent nucleus into two daughters of highly asymmetric masses
- Assuming parent at rest, from the energy conservation $M_{P}c^{2} = M_{D}c^{2} + T_{D} + M_{\alpha}c^{2} + T_{\alpha}$
- Re-organizing the terms, we obtain $T_D + T_\alpha = (M_P - M_D - M_\alpha)c^2 = \Delta M c^2$ ×T

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• Since electron masses cancel, we could use atomic mass expression

$$T_{D} + T_{\alpha} = (M(A,Z) - M(A-4,Z-2) - M(4,2))c^{2} \equiv Q$$

- This is the definition of the <u>disintegration energy</u> or <u>Q</u>-<u>value</u>
 - Difference of rest masses of the initial and final states
 - Q value is equal to the sum of the final state kinetic energies
 - Energy lost during the disintegration process
- For non-relativistic particles, KE are

$$T_D = \frac{1}{2} M_D v_D^2 \qquad T_\alpha = \frac{1}{2} M_\alpha v_\alpha^2$$



- Since the parent is at rest, from the momentum conservation $M_D v_D = M_\alpha v_\alpha \Longrightarrow v_D = \frac{M_\alpha}{M_D} v_\alpha$
- If $M_D = M_{\alpha}$, $v_D = v_{\alpha}$, then $T_D = T_{\alpha}$
- We can write the relationship of KE and Q-value as $T_D + T_\alpha = \frac{1}{2}M_D v_D^2 + \frac{1}{2}M_\alpha v_\alpha^2 = \frac{1}{2}M_D \left(\frac{M_\alpha}{M_D}v_\alpha\right)^2 + \frac{1}{2}M_\alpha v_\alpha^2$ $T_D + T_\alpha = T_\alpha \frac{M_\alpha + M_D}{M_D} \qquad T_\alpha = \frac{M_D}{M_\alpha + M_D}Q$
- This means that T_{α} is unique for the given nuclei
- Direct consequence of 2-body decay of a parent at rest

- KE of the emitted $\boldsymbol{\alpha}$ must be positive
- Thus for an α -decay to occur, it must be an exorthermic process $\Delta M \ge 0, \ Q \ge 0$
- For a massive nuclei, the daughter's KE is

$$T_D = Q - T_\alpha = \frac{M_\alpha}{M_\alpha + M_D} Q = \frac{M_\alpha}{M_D} T_\alpha \ll T_\alpha$$

• Since $M_{\alpha}/M_{D} \approx 4/(A-4)$, we obtain

$$T_{\alpha} \approx \frac{A-4}{A}Q \qquad T_D \approx \frac{4}{A}Q$$



- Most energetic α -particles produced alone
 - Parent nucleus decays to the ground state of a daughter and produces an $\alpha\text{-particle}$ whose KE is the entire Q value
- Less energetic ones accompany photons mostly delayed…
 - Indicates quantum energy levels
 - Parent decays to an excited state of the daughter after emitting an $\boldsymbol{\alpha}$

$${}^{A}X^{Z} \rightarrow {}^{A-4}Y^{*Z-2} + {}^{4}He^{2}$$

Daughter then subsequently de-excite by States of emitting a photon

$$^{A-4}Y^{*Z-2} \rightarrow ^{A-4}Y^{Z-2} + \gamma$$

 Difference in the two Q values corresponds to the energy of the photon

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Nuclear Radiation: α-Decay Example • ²⁴⁰Pu⁹⁴ decay reaction is

$$^{240}Pu^{94} \rightarrow ^{236}U^{92} + ^{4}He^{2}$$

- α particles are observed with 5.17MeV and 5.12 MeV
- Since $Q = \frac{A}{A-4}T_{\alpha}$
- We obtain the two Q-values $Q_1 \approx \frac{240}{236} 5.17 MeV = 5.26 MeV$ $Q_2 \approx \frac{240}{236} 5.12 MeV = 5.21 MeV$
- Which yields photon energy of $E_{\gamma} = \Delta Q = Q_1 Q_2 = 0.05 MeV$
- Consistent with experimental measurement, 45KeV
- Indicates the energy level spacing of order 100KeV for nuclei
 - Compares to order 1eV spacing in atomic levels



Discovery of Alpha and Beta Radiations

- After Bacquerel's discovery of uranium effect on photo-films
 in 1896
- The Curies began study of radio activity in 1898
- Rutherford also studied using a more systematic experimental equipments in 1898
 - Measured the currents created by the radiations
- While Bacquerel concluded that rays are observed in different levels
- Rutherford made the observation using electrometer and determined that there are at least two detectable rays
 - Named α and β rays





Nuclear Radiation: β-Decays

- Three kinds of β -decays
 - Electron emission
 - Nucleus with large N_n
 - Atomic number increases by one
 - Nucleon number stays the same
 - Positron emission
 - Nucleus with many protons
 - Atomic number decreases by one
 - Nucleon number stays the same
 - You can treat nuclear reaction equations algebraically
 - The reaction is valid in the opposite direction as well
 - Any particle moved over the arrow becomes its anti particle

 $^{A}X^{Z} \rightarrow ^{A}Y^{Z+1} + e^{-}$

 $^{A}X^{Z} \rightarrow ^{A}Y^{Z-1} + e^{+}$

Nuclear Radiation: β-Decays

– Electron capture

- Nucleus with many protons
- Absorbs a K-shell atomic electron
- ${}^{A}X^{Z} + e^{-} \rightarrow {}^{A}Y^{Z-1}$
- Proton number decreases by one
- Causes cascade X-ray emission from the transition of remaining atomic electrons
- For β -decay: $\Delta A=0$ and $|\Delta Z|=1$



Nuclear Radiation: β -Decays

- Initially assumed to be 2-body decay ${}_{A}{}_{X^{Z}} \rightarrow {}^{A}{}_{Y^{Z+1}} + e^{-}$
- From energy conservation

$$E_X = E_Y + E_{e^-} = E_Y + T_e + m_e c^2$$

- Since lighter electron carries most the KE $T_e = \left(E_X - E_Y - m_e c^2\right) = \left(m_X - m_Y - m_e\right)c^2 + T_Y = Q - T_Y \approx Q$
- Results in a <u>unique Q value</u> as in α-decay.
- In reality, electrons emitted with continuous E spectrum with an end-point given by the formula above
- Energy conservation is violated!!!! Monday, Oct. 3, 2016 PHYS 3446, F



Nuclear Radiation: β -Decays

- Angular momentum is also in trouble
- In $\beta\text{-decays}$ total number of nucleons is conserved
 - Recall $|\Delta A|=0$ and $|\Delta Z|=1$ in β -decays?
- Electrons are fermions with spin $\frac{1}{2}\hbar$
- Independent of any changes of an integer orbital angular momentum, the total angular momentum cannot be conserved
 - How much does it always differ by? $\frac{1}{2}\hbar$
- Angular momentum conservation is violated!!!



Nuclear Radiation: β-Decays

- In 1931 Pauli postulated an additional particle emitted in βdecays
 - No one observed this particle in experiments
 - Difficult to detect
 - First observation of ν_e in 1956, ν_μ in 1962 and ν_τ in 1977 (direct 2000)
 - Charge is conserved in β -decays
 - Electrically neutral
 - Maximum energy of electrons is the same as the Q value
 - Massless
 - Must conserve the angular momentum
 - Must be a fermion with spin $\frac{1}{2}\hbar$
- This particle is called the neutrino (by Feynman) and expressed as v



Nuclear Radiation: Neutrinos

- Have anti-neutrinos \overline{v} , just like other particles
- Neutrinos and anti-neutrinos are distinguished through the spin projection on momentum
 - Helicity is used to distinguish them $H \propto \vec{p} \cdot \vec{s}$
 - Left-handed (spin and momentum opposite direction) anti-electron-neutrinos are produced in β -decays
 - Right-handed electron-neutrinos are produced in positron emission
 - e⁻ is a particle and e⁺ is the anti-particle to e⁻
 - $-v_e$ is a particle and \overline{v}_e is the anti-particle to v_e



$\beta\text{-}\text{Decay}$ Reaction Equations with Neutrinos

• Electron emission

$${}^{A}X^{Z} \rightarrow {}^{A}Y^{Z+1} + e^{-} + \overline{\nu}_{e}$$

Positron emission

$${}^{A}X^{Z} \rightarrow {}^{A}Y^{Z-1} + e^{+} + V_{e}$$

• Electron capture

$${}^{A}X^{Z} + e^{-} \rightarrow {}^{A}Y^{Z-1} + V_{e}$$



β -Decays with neutrinos

- If the parent nucleus decays from rest, from the conservation of energy $M_p c^2 = T_D + M_D c^2 + T_{\rho^-} + m_e c^2 + T_{\overline{\nu}_e} + m_{\overline{\nu}_e} c^2$
- Thus the Q-value of a β -decay can be written $T_D + T_{e^-} + T_{\overline{v}_e} = \left(M_p - M_D - m_e - m_{\overline{v}_e}\right)c^2 = \Delta Mc^2 = Q$
- Electron emission can only occur if Q>0
- Neglecting all small atomic BE, e emission can occur if

$$Q = \left(M(A,Z) - M(A,Z+1) - m_e - m_{\overline{v}_e} \right) c^2$$

$$\approx \left(M(A,Z) - M(A,Z+1) \right) c^2 \ge 0$$



β -Decays with neutrinos

- Since the daughter nucleus is much heavier than e or $\nu,$ the small recoil energy of daughter can be ignored
 - Thus we can obtain $T_{\rho^-} + T_{\overline{\nu}_a} \approx Q$
- This means that the energy of the electron is not unique and can be any value in the range $0 \le T_{-} \le Q$
 - The maximum electron kinetic energy can be Q
 - This is the reason why the electron energy spectrum is continuous and has an end point (=Q)
- The same can apply to the other two $\beta\text{-decays}$



Particle Numbers

- Baryon numbers: A quantum number assigned to baryons (particles consist of quarks)
 - Mostly conserved in many interactions
 - Baryons: +1
 - Anti-baryons: -1
 - Protons and neutrons are baryons with baryon number +1 each
- Lepton numbers: A quantum number assigned to leptons (electrons, muons, taus and their corresponding neutrinos)
 - Leptons: +1
 - Anti-leptons: -1
 - Must be conserved at all times under SM in each species

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Lepton Numbers

- Three charged leptons exist in nature with their own associated neutrinos $\begin{pmatrix} e^- \\ v_e \end{pmatrix} \begin{pmatrix} \mu^- \\ v_{\mu} \end{pmatrix} \begin{pmatrix} \tau^- \\ v_{\tau} \end{pmatrix}$
- These three types of neutrinos are distinct from each
 other
 - muon neutrinos never produce other leptons than muons or anti-muons

$$\begin{array}{c} \nu_{\mu} + {}^{A}X^{Z} \rightarrow {}^{A}Y^{Z+1} + \mu^{-} \\ \nu_{\mu} + {}^{A}X^{Z} \not \rightarrow {}^{A}Y^{Z+1} + e^{-} \\ \nu_{\mu} + {}^{A}X^{Z} \not \rightarrow {}^{A}Y^{Z+1} + e^{-} \\ \nu_{\mu} + {}^{A}X^{Z} \not \rightarrow {}^{A}Y^{Z+1} + \tau^{-} \\ \end{array}$$
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Lepton Numbers For electron neutrinos

 $\nu_{o} + {}^{A}X^{Z} \rightarrow {}^{A}Y^{Z+1} + e^{-}$ $\nu_e + {}^A X^Z \not\rightarrow {}^A Y^{Z+1} + \mu^ \nu_{a} + {}^{A}X^{Z} \not\rightarrow {}^{A}Y^{Z+1} + \tau^{-}$

For tau neutrinos

 $\nu_{\tau} + {}^{A}X^{Z} \rightarrow {}^{A}Y^{Z+1} + \tau^{-}$ $\nu_{\tau} + {}^{A}X^{Z} \not\rightarrow {}^{A}Y^{Z+1} + e^{-}$ $\nu_{\tau} + {}^{A}X^{Z} \not\rightarrow {}^{A}Y^{Z+1} + \mu^{-}$



Neutrino Mass

• What does neutrino mass do to the β -spectrum?



- The higher end tail shape depends on the mass of the neutrino
 - $-\beta$ -spectrum could be used to measure the mass of neutrinos
 - Very sensitive to the resolution on the device
 - Most stringent direct limit is m_v <2eV/c²
- Non-zero mass of the neutrino means
 - Neutrino Oscillation: Mixing of neutrino species



• β -decay can be written at the nucleon level as:

 $n \rightarrow p + e^- + \overline{v_e} \qquad p \rightarrow n + e^+ + v_e \qquad p + e^- \rightarrow n + v_e$

- Since neutrons are heavier than protons, they can decay to a proton in a free space
 - On the other hand, protons are lighter than neutrons therefore they can only undergo a β -decay within a nucleus
 - Life time of a neutron is about 900sec
 - This life time is a lot longer than nuclear reaction time scale 10^{-23} s or EM scale 10^{-16} s.
- This means that a β -decay is a nuclear phenomenon that does not involve strong nuclear or EM forces
- Fermi postulated a new weak force responsible for $\beta\text{-}$ decay



- Weak forces are short ranged
 - How do we know this?
 - Occurs in the nuclear domain
 - Weakness of the strength is responsible for long life time seen in $\beta\text{-decays}$
- Nucleus does not contain electrons
 - Electrons in β -decays must come from somewhere else
 - Electrons are emitted without time delay
 - The electron must come at the time of decay just like the alphas from a nuclear disintegration
 - $\ \beta$ -decay can be considered to be induced by the weak force



 The transition probability per unit time, the width, can be calculated from perturbation theory using Fermi's Golden rule

$$P = \frac{2\pi}{\mathsf{h}} \left| H_{fi} \right|^2 \rho \left(E_f \right)$$

Where the weak interaction Hamiltonian is

$$H_{fi} = \left\langle f \left| H_{wk} \right| i \right\rangle = \int d^3 x \psi_f^* \left(x \right) H_{wk} \psi_i \left(x \right)$$



- Based on β -decay reaction equations, the H_{wk} must be a four fermionic states
 - H_{wk} proposed by Fermi in 1933 is a <u>four-fermion interaction</u> or current-current interaction
 - Relativistic
 - Agreed rather well with experiments for low energy β -decays
- Parity violation
 - There are only left-handed neutrinos and right-handed anti-neutrinos
 - A system is parity invariant if it does not change under inversion of spatial coordinates
 - The spin $r \to -r$, $p \to -p$ $\Rightarrow L = r \times p = (-r) \times (-p) = L$
 - The handedness, helicity, changes upon the spatial inversion since the direction of the motion changes while the spin direction does not
 - Since there is no right handed neutrinos, parity must be violated in weak interactions

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Gamma Decays

- When a heavy nuclei undergo alpha and beta decays, the daughters get into an excited state
 - Must either break apart
 - Or emit another particle
 - To bring the daughter into its ground state



- Typical energies of photons in $\gamma\text{-decays}$ are a few MeV's
 - These decays are EM interactions thus the life time is on the order of 10⁻¹⁶sec.
- Photons carry one unit of angular momentum
 - Parity is conserved in this decay



Assignments

- 1. Reading assignment: CH 4.3
- 2. End of the chapter problems: 3.2
- 3. Derive the following equations:
 - Eq. 4.8 starting from conservation of energy
 - Eq. 4.11 both the formula
- Due for these homework problems is next Monday, Oct. 16.

