## PHYS 5326 – Lecture #4

Wednesday, Jan. 31, 2007 Dr. Jae Yu

- 1. QCD Evolution of PDF
- 2. Measurement of  $Sin^2\theta_W$
- 3. Formalism of  $Sin^2\theta_W$  in v-N DIS
- 4. Improvements in  $Sin^2\theta_W$
- 5. Interpretation of  $\text{Sin}^2\theta_{\text{W}}$  and Its Link to Higgs





## **DGLAP QCD Evolution Equations**

 The evolution equations by Dokshitzer-Gribov-Lipatov-Altarelli-Parisi provide mechanism to evolve PDF's to any kinematic regime or momentum scale

$$\frac{dq^{NS}(x,M^{2})}{d\ln M^{2}} = \sum_{i} q^{i} - \overline{q}^{i} = u_{V} + d_{V} = \frac{\alpha_{s}(\mu^{2})}{2\pi} \int_{x}^{1} \frac{dy}{y} q^{NS}(y,M^{2}) P_{qq}\left(\frac{x}{y}\right)$$
$$\frac{dq^{S}(x,M^{2})}{d\ln M^{2}} = \sum_{i} q^{i} + \overline{q}^{i} = \frac{\alpha_{s}(\mu^{2})}{2\pi} \int_{x}^{1} \frac{dy}{y} \left[ q^{NS}(y,M^{2}) P_{qq}^{s}\left(\frac{x}{y}\right) + G(y,M^{2}) P_{qG}^{s}\left(\frac{x}{y}\right) \right]$$
$$\frac{dG(x,M^{2})}{d\ln M^{2}} = \frac{\alpha_{s}(\mu^{2})}{2\pi} \int_{x}^{1} \frac{dy}{y} \left[ q^{S}(y,M^{2}) P_{Gq}^{s}\left(\frac{x}{y}\right) + G(y,M^{2}) P_{GG}\left(\frac{x}{y}\right) \right]$$

 $P_{ij}(x/y)$ : Splitting function which describes the probability for a parton *i* with momentum y get resolved as a parton *j* with momentum x<y

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# Feynman Diagrams for Parton Splitting

**LO: Ο(**α<sub>s</sub>)

#### NLO: $O(\alpha_s^2)$





### **Electroweak Theory**

- Standard Model unifies Weak and EM to SU(2)xU(1) gauge theory
  - Weak neutral current interaction
  - Measured physical parameters related to mixing parameters for the couplings

$$g' = g \tan \theta_W \quad e = g \sin \theta_W \quad G_F = \frac{g^2 \sqrt{2}}{8M_W^2} \quad \frac{M_W}{M_Z} = \cos \theta_W$$

- Neutrinos in this picture are unique because they only interact through left-handed weak interactions → Probe weak sector only
  - Less complication in some measurements, such as proton structure



## $\text{sin}^2\theta_{\text{W}}$ and $\nu\text{-N}$ scattering

- In the electroweak sector of the Standard Model, it is not known *a priori* what the mixture of electrically neutral electromagnetic and weak mediator is -> This fractional mixture is given by the mixing angle
- Within the on-shell renormalization scheme,  $sin^2\theta_W$  is:

$$\sin^2 \theta_w^{On-Shell} = 1 - \frac{M_W^2}{\rho_0 M_Z^2}$$

Provides independent measurement of M<sub>W</sub> & information to pin down M<sub>Higgs</sub> via higher order loop corrections, in comparable uncertainty to direct measurements
 Measures light quark couplings → Sensitive to other types (anomalous) of couplings

•In other words, sensitive to physics beyond SM  $\rightarrow$  New vector bosons, compositeness, v-oscillations, etc



## **EW Higher Order Corrections**

- LO GSW requires three parameters:  $\alpha$ , G<sub>F</sub> and M<sub>Z</sub>
- Higher order corrections bring in dependences to two additional parameters:  $M_{Top}$  and  $M_{Higgs}$





• Cross section ratios between NC and CC proportional to  $sin^2\theta_w$ 

• Llewellyn Smith Formula:

$$\mathbf{R}^{\nu(\overline{\nu})} = \frac{\sigma_{\mathrm{NC}}^{\nu(\overline{\nu})}}{\sigma_{\mathrm{CC}}^{\nu(\overline{\nu})}} = \rho^{2} \left( \frac{1}{2} - \sin^{2}\theta_{\mathrm{W}} + \frac{5}{9}\sin^{4}\theta_{\mathrm{W}} \left( 1 + \frac{\sigma_{\mathrm{CC}}^{\overline{\nu}(\nu)}}{\sigma_{\mathrm{CC}}^{\nu(\overline{\nu})}} \right) \right)$$





- Very small cross section → Heavy neutrino target
- $v_e$  are the killers (CC events look the same as NC events)



## How Can Events be Separated?



#### **Experimental Variable**

#### Define an Experimental Length variable

→ Distinguishes CC from NC experimentally in statistical manner



## Past Experimental Results $sin^2 \theta_W^{On-Shell} = 1 - \frac{M_W^2}{M_Z^2} = 0.2277 \pm 0.0036$ $\Rightarrow M_W^{On-Shell} = 80.14 \pm 0.19 \text{GeV/c}^2$

The yellow band represents a correlated uncertainty!!

## Improvements on Measurements

- Asses the uncertainties from previous measurements
- Determine what the sources of largest theoretical and experimental uncertainties are
- Provide new methods to reduce large uncertainties



# $sin^2\theta_W$ Theoretical Uncertainty

 Significant correlated error from CC production of charm quark (m<sub>c</sub>) modeled by slow rescaling mechanism



• Suggestion by Paschos-Wolfenstein by separating v and  $\overline{v}$  beams:

$$\mathbf{R}^{-} = \frac{\boldsymbol{\sigma}_{NC}^{\nu} - \boldsymbol{\sigma}_{NC}^{\overline{\nu}}}{\boldsymbol{\sigma}_{CC}^{\nu} - \boldsymbol{\sigma}_{CC}^{\overline{\nu}}} = \boldsymbol{\rho}^{2} \left(\frac{1}{2} - \sin^{2}\boldsymbol{\theta}_{W}\right) = \frac{\mathbf{R}^{\nu} - \mathbf{R}^{\overline{\nu}}}{1 - \mathbf{r}}$$

→ Reduce charm CC production error by subtracting sea quark contributions
 → Only valence u, d, and s contributes while sea quark contributions cancel out
 → Massive quark production through Cabbio suppressed d<sub>v</sub> quarks only

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### Improving Experimental Uncertainties

- Electron neutrinos,  $\nu_{\rm e}$ , in the beam fakes NC events from CC interactions
  - If the production cross section is well known, the effect will be smaller but since majority come from neutral K (K<sub>L</sub>) whose x-sec is known only to 20%, this is a source of large experimental uncertainty
- Need to come up with a beamline that separates neutrinos from anti-neutrinos



#### **Event Contamination and Backgrounds**

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•SHORT  $v_{\mu}$  CC's (20% v, 10%  $\overline{v}$ )  $\mu$  exit and rangeout •SHORT  $v_{e}$  CC's (5%)  $v_{e}N \rightarrow eX$ •Cosmic Rays (0.9%)

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•LONG  $v_{\mu}$  NC's (0.7%) hadron shower punch-through effects

•Hard μ Brem(0.2%) Deep μ events





Sources of experimental uncertainties kept small, through modeling using  $\boldsymbol{\nu}$  and TB data

Effect	<mark>Size(</mark> δsin²θ <sub>w</sub> )	Tools	
Z <sub>vert</sub>	0.001/inch	μ⁺μ⁻ events	
X <sub>vert</sub> & Y <sub>vert</sub>	0.001	MC	
Counter Noise	0.00035	TB μ's	
Counter Efficiency	0.0002	v events	
Counter active area	0.0025/inch	ν CC, TB	
Hadron shower length	0.0015/cntr	TB $\pi$ 's and k's	
Energy scale	0.001/1%	ТВ	
Muon Energy Deposit	0.004	v CC	
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#### Measurements of $v_e$ Flux

- Use well known processes (Ke3:  $\mathbf{K}^{\pm} \rightarrow \pi^{0} e^{\pm} \overset{(-)}{\nu}_{e}$ ) ۲
- Shower Shape Analysis can provide direct measurement  $v_e$  events, though less precise



 $N_{meas}/N_{MC}$ **1.05**  $\pm$  **0.03**  $(\nu_{e})$  used for  $\nu_{e}$   $\Rightarrow \delta R_{v}^{exp} \sim 0.0005$  $1.01 \pm 0.04 \left( \overline{\nu}_{e} \right)$ 

Weighted average

- $v_e$  from very short events (E<sub>v</sub>>180 GeV)
  - Precise measurement of  $v_e$  flux in the tail region of flux  $\rightarrow$  ~35% more  $\overline{v}_{e}$  in  $\overline{v}$  than predicted
  - Had to require ( $E_{had}$ <180 GeV) due to ADC saturation

Results in  $\sin^2\theta_w$  shifts by +0.002





# MC to Relate $R_v^{exp}$ to $R^v$ and $sin^2\theta_W$

- Parton Distribution Model
  - − Correct for details of PDF model → Used CCFR data for PDF
  - Model cross over from short  $\nu_{\mu}$  CC events Neutrino xsec vs y at 190 GeV Antineutrino xsec vs y at 190 GeV 324 v=1.019 CFR Data 02 04 05 05 30 3.0 +0 0.2 0.4 0.0 C.B d'occide c=0.05 <=0.125 s=0.125 1.00 205 0.4 0.8 7.5 x-3,175 -0.175 <-0.225 50.5 02 0.8 0.7 0.7
- Neutrino Fluxes
  - $v_{\mu}, v_{e}, \overline{v}_{\mu}, \overline{v}_{e}$  in the two running modes
  - $v_e$  CC events always look short
- Shower length modeling
  - Correct for short events that look long
- Detector response vs energy, position, and time
  - Continuous testbeam running minimizes systematics

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#### $sin^2\theta_W\, Fit \ to \ R_\nu^{\ exp} \ and \ R \ _\nu^{\ exp}$

- Thanks to the separate beam  $\rightarrow$  Measure R<sup>v</sup>'s separately
- Use MC to simultaneously fit  $R_{\nu}^{exp}$  and  $R_{\overline{\nu}}^{exp}$  to  $sin^2\theta_W$  and  $m_c$ , and  $sin^2\theta_W$  and  $\rho$

$$\mathsf{R}^{\nu(\overline{\nu})} = \frac{\sigma_{\mathsf{NC}}^{\nu(\overline{\nu})}}{\sigma_{\mathsf{CC}}^{\nu(\overline{\nu})}} = \rho^{2} \left(\frac{1}{2} - \sin^{2}\theta_{\mathsf{W}} + \frac{5}{9}\sin^{4}\theta_{\mathsf{W}} \left(1 + \frac{\sigma_{\mathsf{CC}}^{\overline{\nu}(\nu)}}{\sigma_{\mathsf{CC}}^{\nu(\overline{\nu})}}\right)\right)$$

- R<sup>v</sup> Sensitive to sin<sup>2</sup> $\theta_W$  while R  $\bar{v}$  isn't, so R<sup>v</sup> is used to extract sin<sup>2</sup> $\theta_W$  and R  $\bar{v}$  to control systematics
- Single parameter fit, using SM values for EW parameters ( $\rho_0$ =1)

 $\sin^2 \theta_w = 0.2277 \pm 0.0013 \text{ (stat)} \pm 0.0009 \text{ (syst)}$ 

m<sub>c</sub> = 1.32  $\pm$  0.09 (stat)  $\pm$  0.06 (syst) w/m<sub>c</sub> = 1.38  $\pm$  0.14 GeV/c<sup>2</sup> as input



#### NuTeV sin<sup>2</sup> $\theta_{W}$ Uncertainties

	V V		
Source of Uncertainty	δ sin²θ <sub>w</sub>	Dominant	
Statistical	0.00135	uncertainty	
$v_{e}$ flux	0.00039	1-Loop Electroweak Radiative Corrections based on Bardin, Dokuchaeva JINR-E2-86-2 60 (1986)	
Event Length	0.00046		
Energy Measurements	0.00018		
Total Experimental Systematics	0.00063		
CC Charm production, sea quarks	0.00047		
Higher Twist	0.00014		
Non-isoscalar target	0.00005	δsin <sup>2</sup> θ <sup>(On-shell)</sup> <sub>w</sub> = -0.00022 × $\left(\frac{M_t^2 - (175 \text{GeV})^2}{2}\right)^2$	
$\sigma^{\overline{ u}}/\sigma^{ u}$	0.00022	$(50 \text{GeV})^2$	
RadiativeCorrection	0.00011	$+0.00032 \times \ln\left(\frac{M_{H}}{10000000000000000000000000000000000$	
R <sub>L</sub>	0.00032	(150GeV)	
Total Physics Model Systmatics	0.00064		
Total Systematic Uncertainty	0.00162		
∆M <sub>w</sub> (GeV/c²)	0.08		
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#### NuTeV vs CCFR Uncertainty Comparisons





#### Comparison of New $sin^2\theta_W$

$$\begin{split} \sin^2\theta_W^{\text{On-Shell}} &= 0.2277 \pm 0.0013 \text{ (stat)} \pm 0.0009 \text{ (syst)} \\ &\quad \sin^2\theta_W^{\text{On-Shell}} = 1 - \frac{M_W^2}{M_Z^2} \\ &\implies M_W^{\text{On-Shell}} = 80.14 \pm 0.08 \text{ GeV/c}^2 \end{split}$$

Comparable precision but value smaller than other measurements W-Boson Mass [GeV]



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