

PHYS 3313 – Section 001

Lecture #12

Monday, Feb. 27, 2017

*Dr. **Jaehoon** **Yu***

- Photoelectric Effect
- Compton Effect
- Pair production/Pair annihilation
- Rutherford Scattering Experiment and Rutherford Atomic Model



Announcements

- Midterm Exam

- In class Wednesday, March. 8
- Covers from CH1.1 through what we learn March 6 plus the math refresher in the appendices
- Mid-term exam constitutes 20% of the total
- **Please do NOT miss the exam! You will get an F if you miss it.**
- BYOF: You may bring a one 8.5x11.5 sheet (front and back) of handwritten formulae and values of constants for the exam
- No derivations, word definitions or solutions of any problems!
 - Eg., Lorentz velocity transformation NOT allowed!
- No additional formulae or values of constants will be provided!

- Quiz 2 results

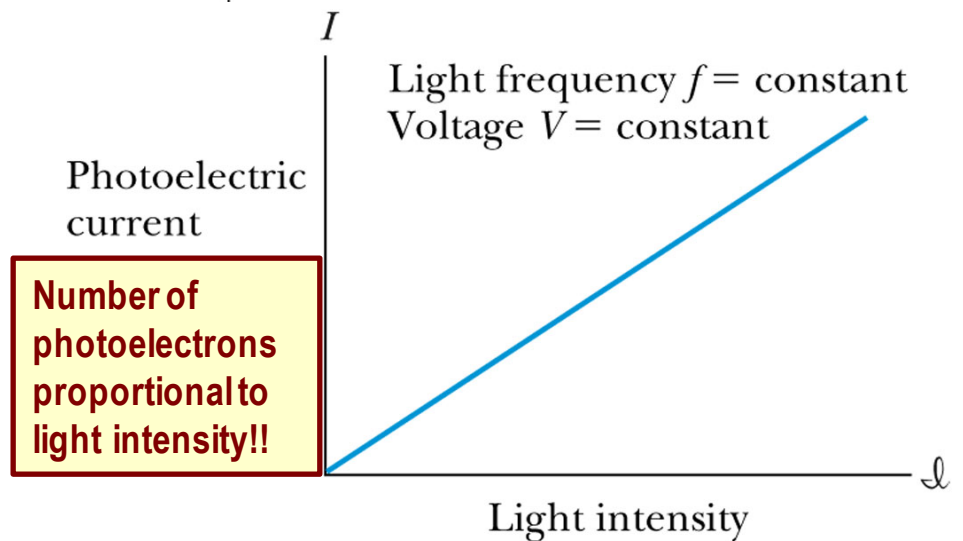
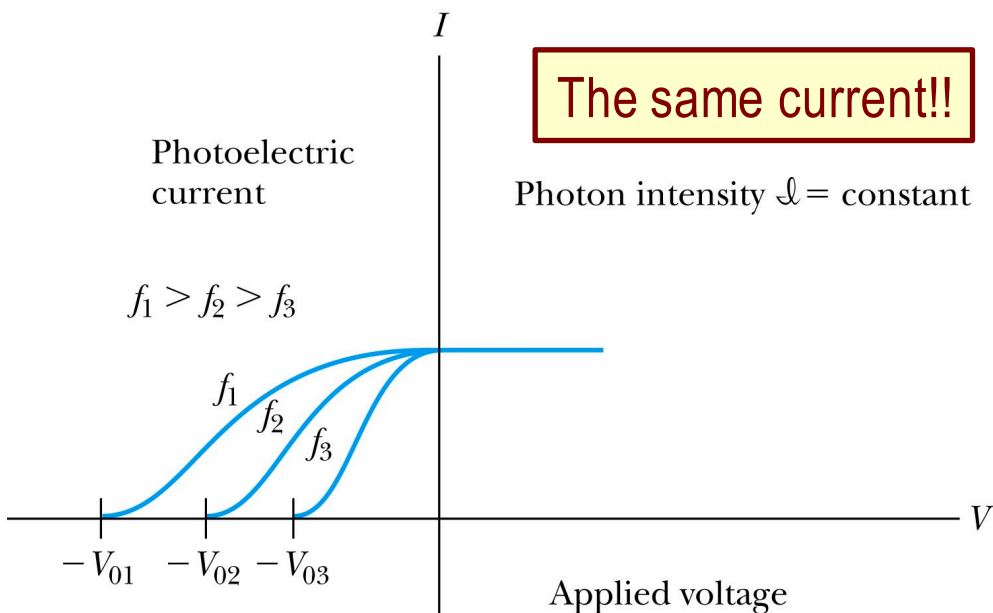
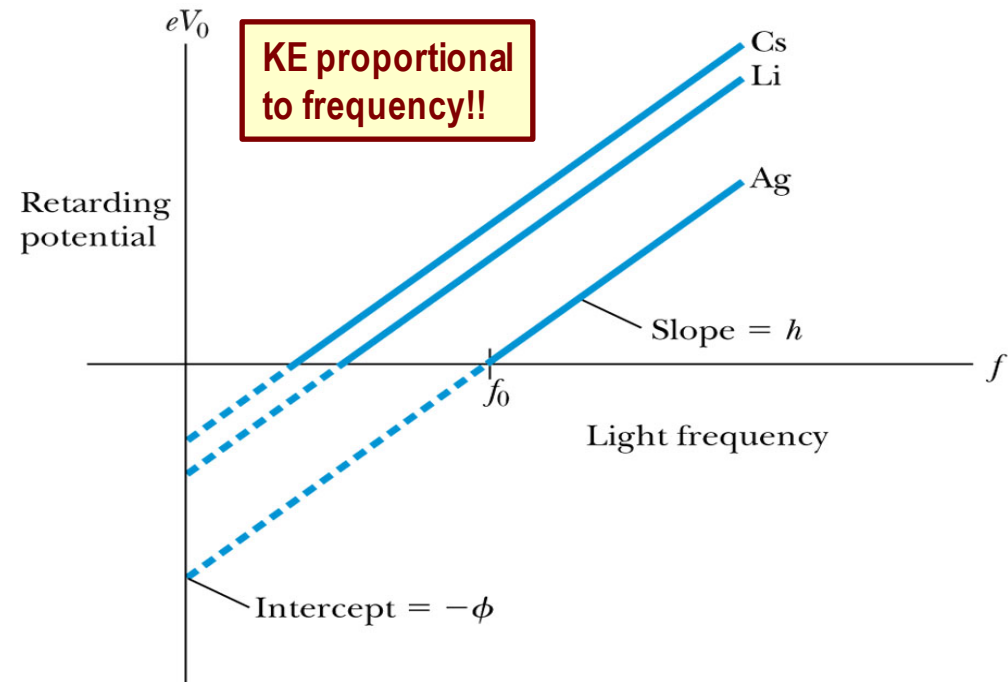
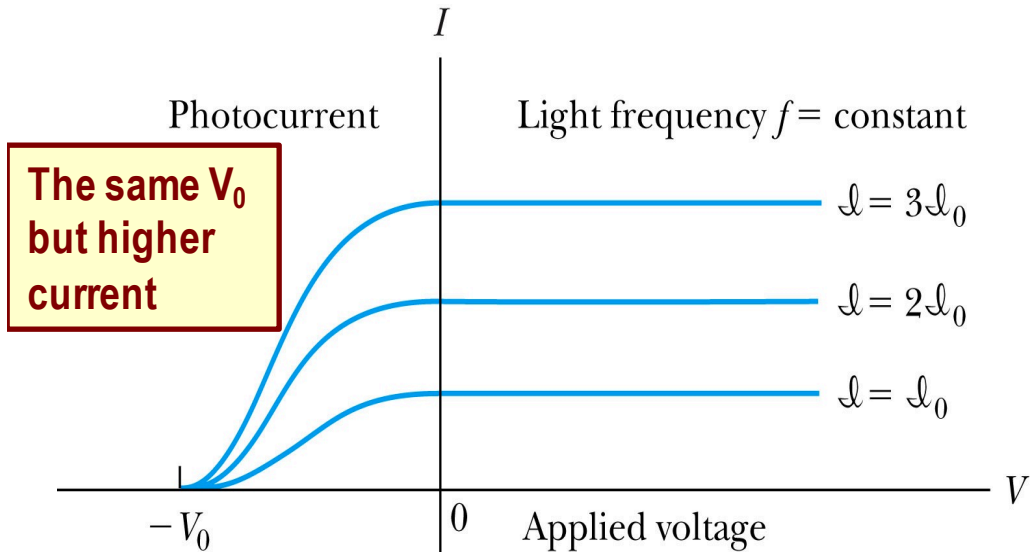
- Class Average: 34.5/60
- Equivalent to : 57.5/100
 - Previous result: 20.8/100
- Class top score: 60/60

Monday, Feb. 27, 2017



PHYS 3313-001, Spring 2017
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Experimental Observations



Wednesday, Feb. 22,
2017



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Dr. Jaehoon Yu

Summary of Experimental Observations

- Light intensity does not affect the KE of the photoelectrons
- The max KE of the photoelectrons for a given emitter material depends only on the frequency of the light
- The smaller the work function ϕ of the emitter material, the smaller is the threshold frequency of the light that can eject photoelectrons.
- When the photoelectrons are produced, their number is proportional to the intensity of light.
- The photoelectrons are emitted almost instantly following the illumination of the photocathode, independent of the intensity of the light. → Totally unexplained by classical physics

Einstein's Theory of Photoelectric Effect

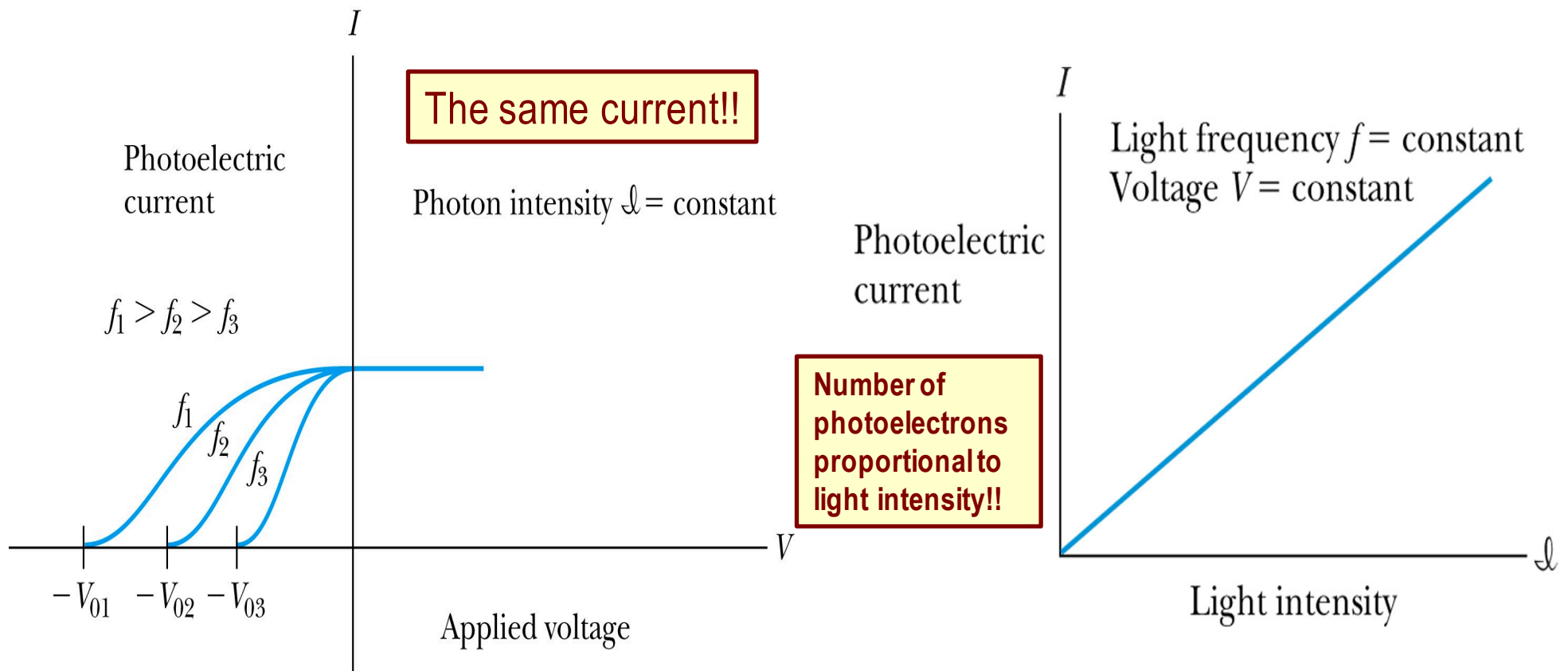
- Einstein suggested that the electromagnetic radiation of the light is quantized into particles called **photons**. Each photon has the energy quantum:

$$E = hf$$

- where f is the frequency of the light and h is Planck's constant.
- The photon travels at the speed of light in vacuum, and its wavelength is given by

$$\lambda f = c$$

Experimental Observations



Einstein's Theory

- Conservation of energy yields:

Energy Before(photon)=Energy After (electron)

$$hf = \phi + KE(\text{photoelectron})$$

where ϕ is the work function of the metal

The photon energy can then be written

$$hf = \phi + \frac{1}{2}mv_{\text{max}}^2$$

- The retarding potentials measure the KE of the most energetic photoelectrons.

$$eV_0 = \frac{1}{2}mv_{\text{max}}^2$$

Quantum Interpretation

- KE of the electron depends only on the light frequency and the work function ϕ of the material not the light intensity at all

$$\frac{1}{2}mv_{\max}^2 = eV_0 = hf - \phi$$

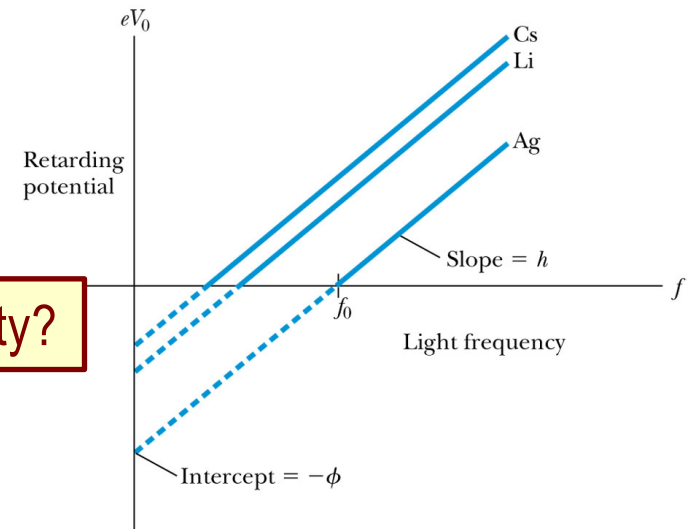
- Einstein in 1905 predicted that the stopping potential was linearly proportional to the light frequency, with the slope h , the same constant found by Planck.

$$eV_0 = \frac{1}{2}mv_{\max}^2 = hf - hf_0 = h(f - f_0)$$

- From this, Einstein concluded that light is a particle with energy:

$$E = hf = \frac{hc}{\lambda}$$

Was he already thinking about particle/wave duality?



Ex 3.11: Photoelectric Effect

- Light of wavelength 400nm is incident upon lithium ($\phi=2.93\text{eV}$). Calculate (a) the photon energy (eV) and (b) the stopping potential V_0 .
- Since the wavelength is known, we use plank's formula:

$$E = hf = \frac{hc}{\lambda} = \frac{(1.626 \times 10^{-34} \text{ J} \cdot \text{s})(3 \times 10^8 \text{ m/s})}{400 \times 10^{-9} \text{ m}} = 3.10 \text{ eV}$$

- The stopping potential can be obtained using Einstein's formula for photoelectron energy

$$eV_0 = hf - \phi = E - \phi$$

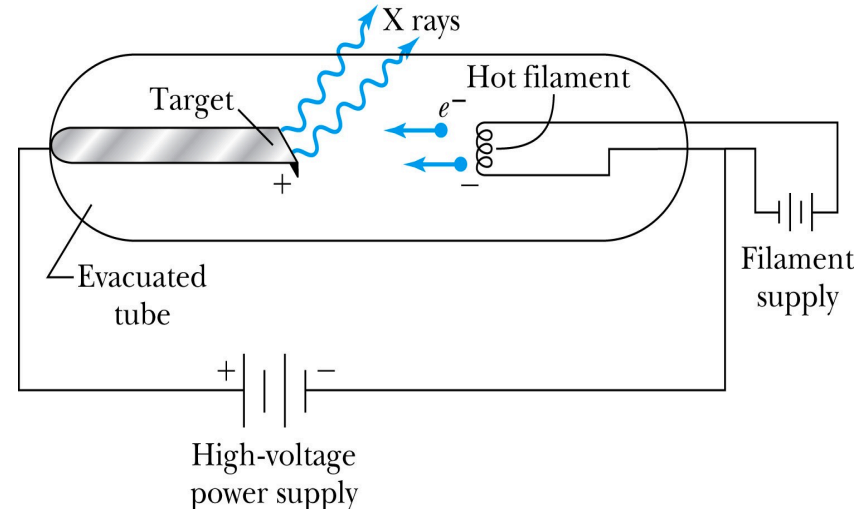
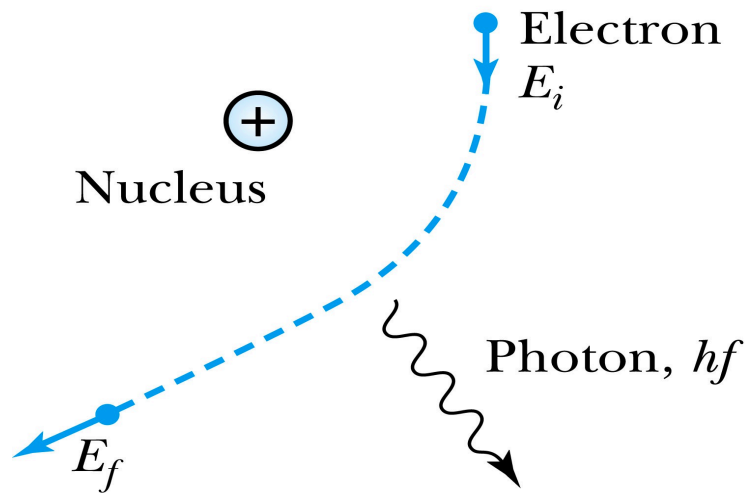
$$V_0 = \frac{E - \phi}{e} = \frac{(3.10 - 2.93) \text{ eV}}{e} = 0.17 \text{ V}$$

X-Ray Production

- **Bremsstrahlung** (German word for braking radiation): Radiation of a photon from an energetic electron passing through matter due to an acceleration
- Since linear momentum must be conserved, the nucleus absorbs very little energy, and it is ignored. The final energy of the electron is determined from the conservation of energy

$$E_f = E_i - hf$$

- An electron that loses a large amount of energy will produce an X-ray photon.
 - Current passing through a filament produces copious numbers of electrons by thermionic emission.
 - These electrons are focused by the cathode structure into a beam and are accelerated by potential differences of thousands of volts until they impinge on a metal anode surface, producing x rays by bremsstrahlung as they stop in the anode material
 - X-ray wavelengths range 0.01 – 10nm. What is the minimum energy of an electron to produce X-ray?



Inverse Photoelectric Effect.

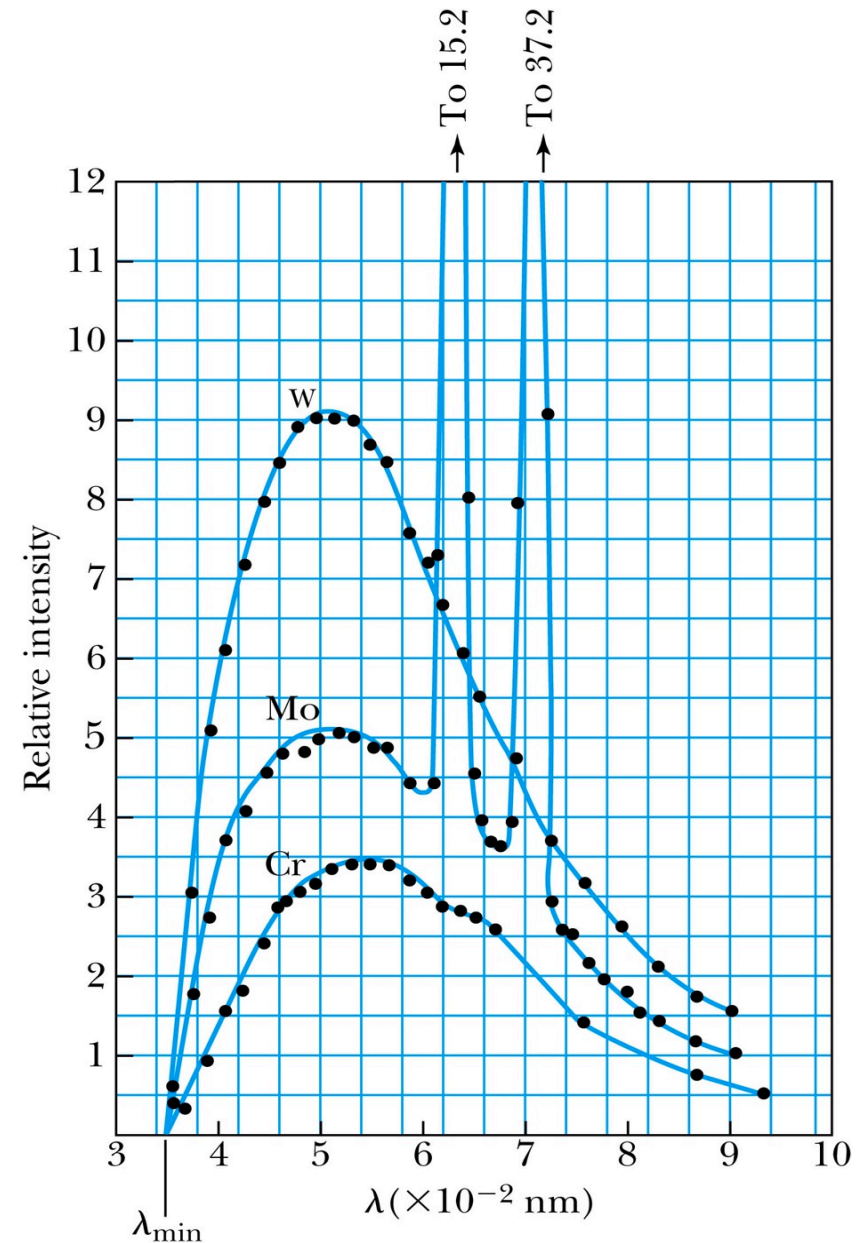
- Conservation of energy requires that the electron KE equals the maximum photon energy

- Work function neglected since it's small compared to the potential energy of the electron.

- This is the **Duane-Hunt limit**

- The photon wavelength depends only on the accelerating voltage
- The same for all targets.

$$eV_0 = hf_{\max} = \frac{hc}{\lambda_{\min}}$$
$$\lambda_{\min} = \frac{hc}{eV_0} = \frac{1.24 \times 10^{-6} V \cdot m}{V_0}$$



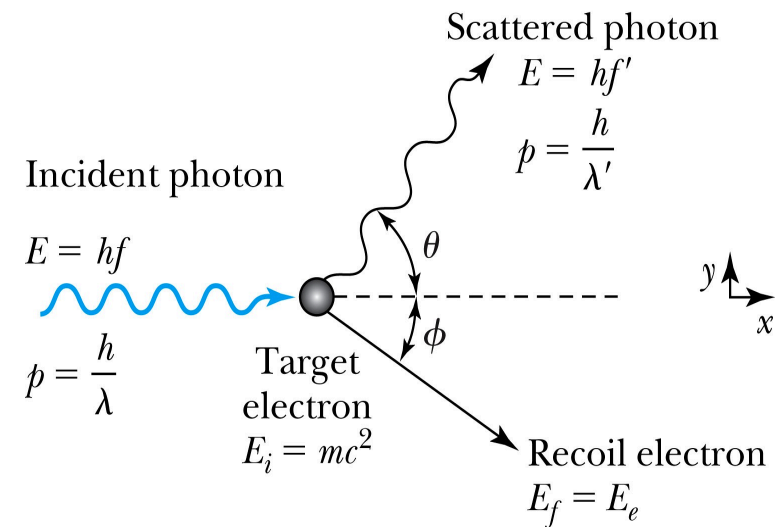
Compton Effect

- When a photon enters matter, it is likely to interact with one of the atomic electrons.
- The photon is scattered from only one electron
- The laws of conservation of energy and momentum apply as in any elastic collision between two particles. The momentum of a particle moving at the speed of light is

$$p = \frac{E}{c} = \frac{hf}{c} = \frac{h}{\lambda}$$

- The electron energy can be written as

$$E_e^2 = (m_e c^2)^2 + p_e^2 c^2$$



- Change of the scattered photon wavelength is known as the **Compton effect**:

$$\Delta\lambda = \lambda' - \lambda = \frac{h}{m_e c} (1 - \cos\theta)$$

Pair Production and Annihilation

- If a photon can create an electron, it must also create a positive charge to balance charge conservation.
- In 1932, C. D. Anderson observed a positively charged electron (e^+) in cosmic radiation. This particle, called a positron, had been predicted to exist several years earlier by P. A. M. Dirac.
- A photon's energy can be converted entirely into an electron and a positron in a process called **pair production**.
 - Can only happen inside a material
 - How much energy do you think is needed?

$$\gamma \rightarrow e^- + e^+$$

Pair Production in Empty Space?

- Energy conservation for pair production in empty space

$$hf = E_+ + E_- + K.E.$$

- Momentum conservation yields

$$hf = p_- c \cos \theta + p_+ c \cos \theta$$

- Thus max momentum exchange $hf_{\max} = p_- c + p_+ c$

- Recall that the total energy for a particle can be written as

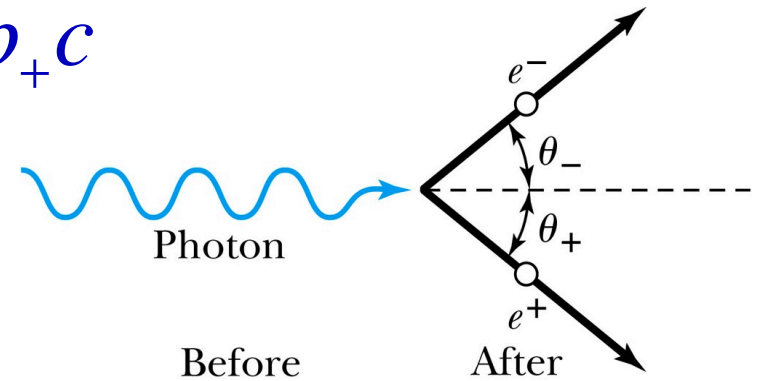
$$E_{\pm}^2 = p_{\pm}^2 c^2 + m_e^2 c^4$$

However this yields a contradiction: $hf > p_- c + p_+ c$

and hence the conversion of energy in empty space is impossible and thus pair production cannot happen in empty space

Pair Production in Matter

- Since the relations $hf_{\max} = p_-c + p_+c$ and $hf > p_-c + p_+c$ contradict each other, a photon can not produce an electron and a positron in empty space.



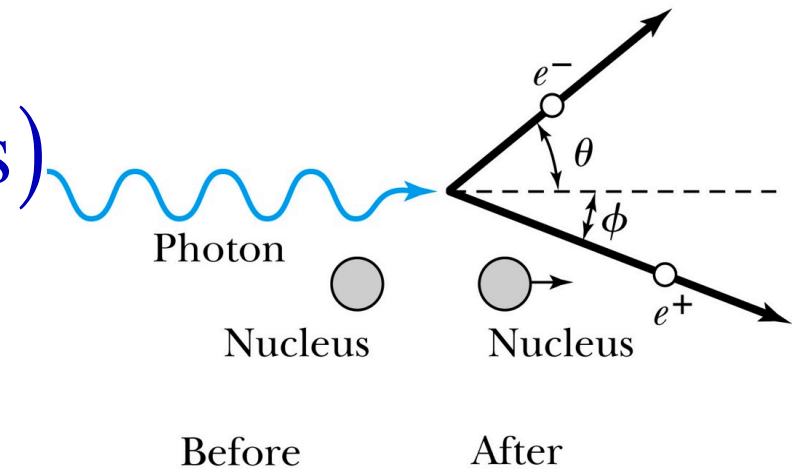
(a) Free space (**cannot occur**)

- In the presence of matter, however, the nucleus absorbs some energy and momentum.

$$hf = E_- + E_+ + K.E.(\text{nucleus})$$

- The photon energy required for pair production in the presence of matter is

$$hf > 2m_e c^2$$



(b) Beside nucleus

Pair Annihilation

- A positron going through matter will likely **annihilate** with an electron.
- A positron is drawn to an electron and form an atom-like configuration called **positronium**.
- Pair annihilation in empty space will produce two photons to conserve momentum. Annihilation near a nucleus can result in a single photon.

- Conservation of energy: $2m_e c^2 \approx hf_1 + hf_2$

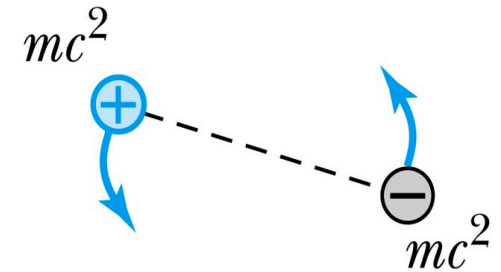
- Conservation of momentum: $0 = \frac{hf_1}{c} - \frac{hf_2}{c}$

- The two photons will be almost identical, so that

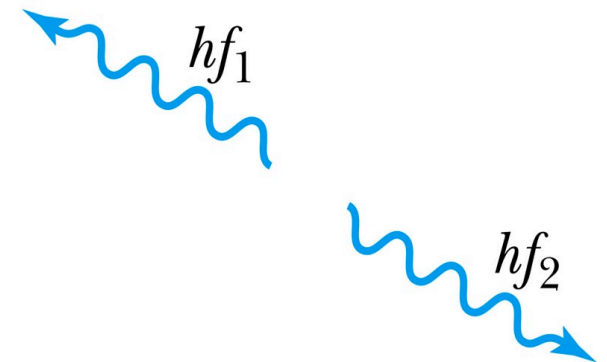
$$f_1 = f_2 = f$$

- The two photons from a positronium annihilation will move in the opposite directions with an energy of:

$$hf = m_e c^2 = 0.511 \text{ MeV}$$



(a) Positronium,
before decay
(schematic only)



(b) After annihilation