PHYS 1443 – Section 001 Lecture #1

Tuesday, May 29, 2007 Dr. Jaehoon Yu

- Who am I?
- How is this class organized?
- What is Physics?
- What do we want from this class?
- Brief history of physics
- Standards and units
- Dimensional Analysis
- Fundamentals
- One Dimensional Motion

Today's homework is homework #1, due 7pm, this Friday!!

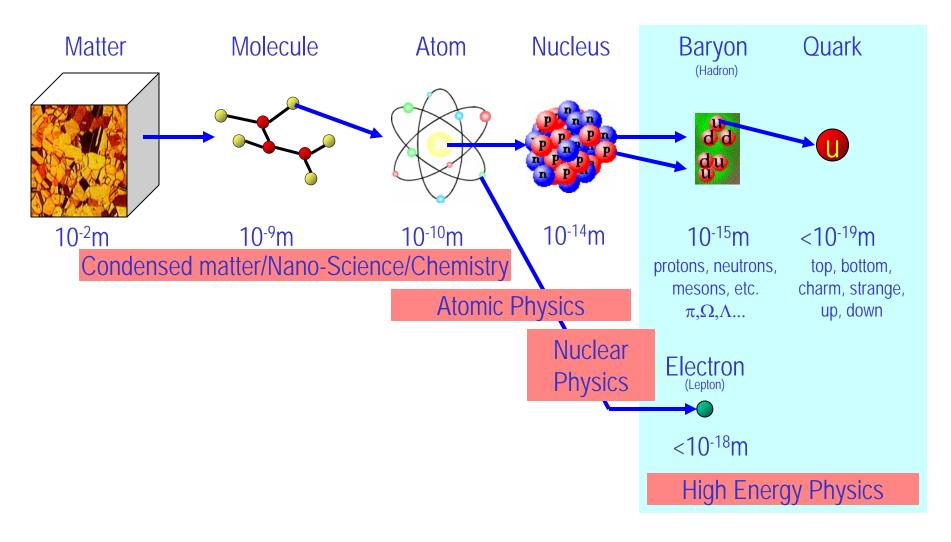
Announcements

- Reading assignment #1: Read and follow through all sections in appendices A and B by Thursday, May 31
 - There will be a quiz on Thursday, May 31, on this reading assignment
- Wednesday and Thursday classes will be taught by Mr. V. Kaushik
 - Tomorrow's class will be slightly shorter

Who am I?

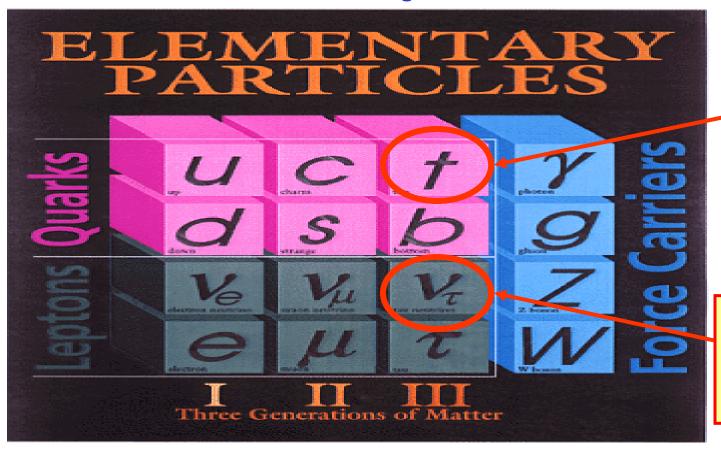
- Name: Dr. Jaehoon Yu (You can call me <u>Dr. Yu</u>)
- Office: Rm 342, Chemistry and Physics Building
- Extension: x22814, E-mail: <u>jaehoonyu@uta.edu</u>
- My profession: High Energy Physics (HEP)
 - Collide particles (protons on anti-protons or electrons on anti-electrons, positrons) at the energies equivalent to 10,000 Trillion degrees
 - To understand
 - Fundamental constituents of matter
 - Interactions or forces between the constituents
 - Origin of Mass
 - Creation of Universe (Big Bang Theory)
 - A pure scientific research activity
 - Direct use of the fundamental laws we find may take longer than we want but
 - Indirect product of research contribute to every day lives; eg. WWW

Structure of Matter



The Standard Model

Assumes the following fundamental structure:



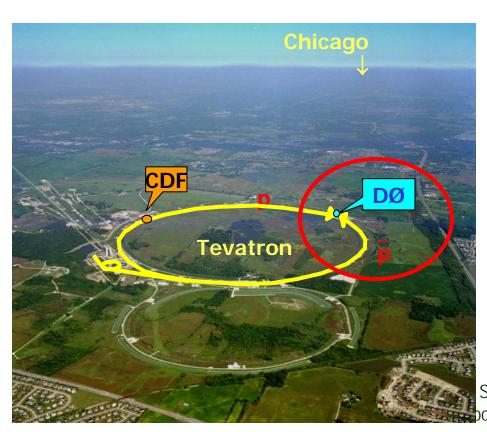
Discovered in 1995

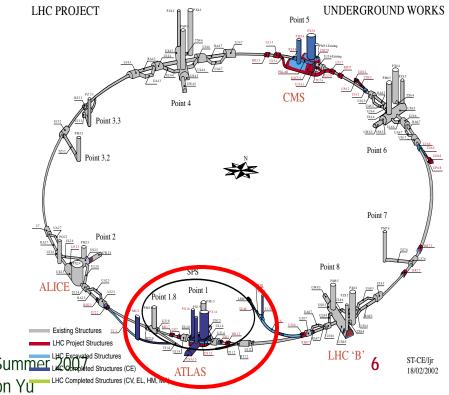
Directly observed in 2000

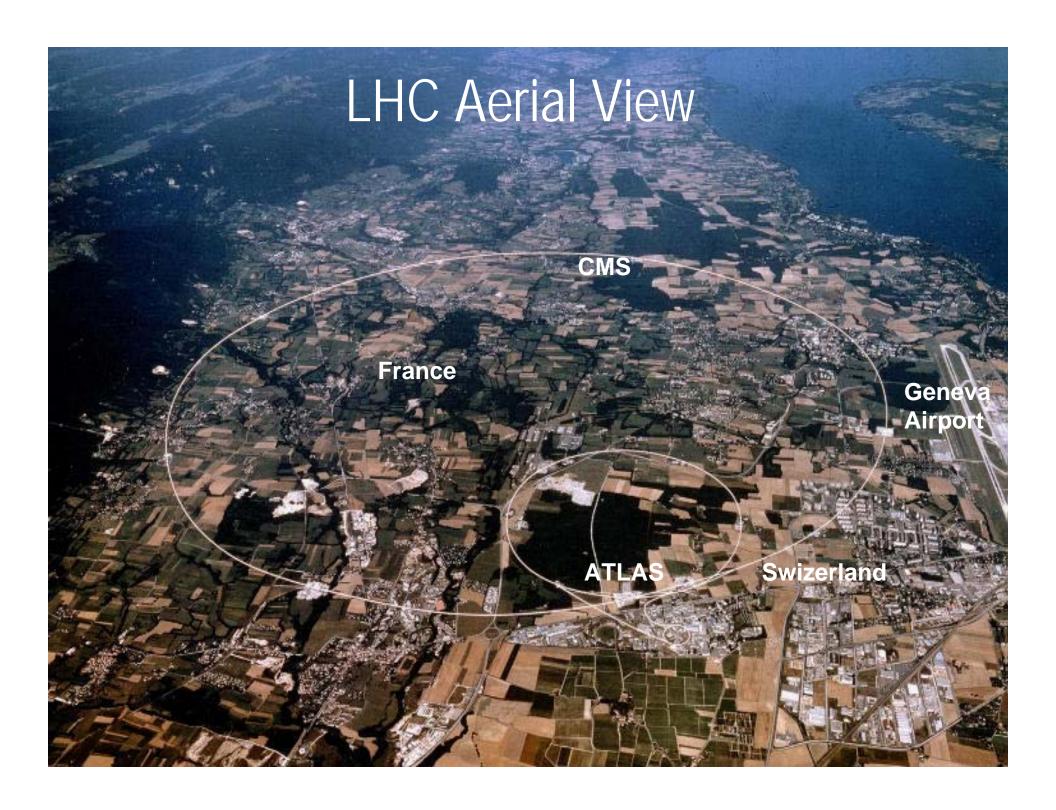
Fermilab Tevatron and LHC at CERN

- Present world's Highest Energy protonanti-proton collider
 - 4km circumference
 - E_{cm} =1.96 TeV (=6.3x10⁻⁷J/p→ 13M Joules on 10⁻⁴m²)
 - ⇒ Equivalent to the kinetic energy of a 20t truck at a speed 81mi/hr → 130km/hr

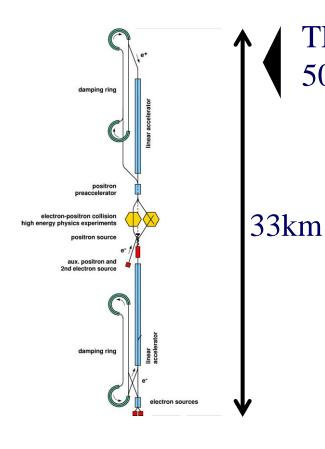
- World's Highest Energy proton-proton collider end of this year
 - 27km circumference
 - E_{cm} =14 TeV (=44x10⁻⁷J/p→ 1000M Joules on 10⁻⁴m²)
 - ⇒ Equivalent to the kinetic energy of a 20t truck at a speed 711mi/hr→1140km/hr







The International Linear Collider



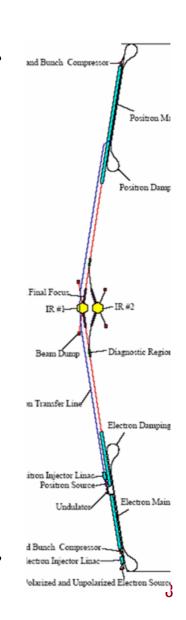
TESLA TDR 500 GeV (800 GeV)

- Long~ linear electronposition colliders
- 10 15 years from now
- Takes 10 years to build an accelerator and the detectors

47 km

US Options Study 500 GeV (1 TeV)





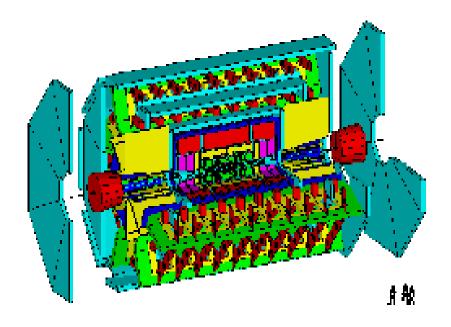
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DØ Detector

Central Tracking Calorimeter Muon Tracking 50',

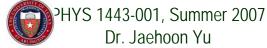
- Weighs 5000 tons and 5 story tall
- Can inspect 3,000,000 collisions/second
- Record 75 collisions/second
- Records approximately 10,000,000 bytes/second
- Records 0.5x10¹⁵ (500,000,000,000,000) bytes per year (0.5 PetaBytes).

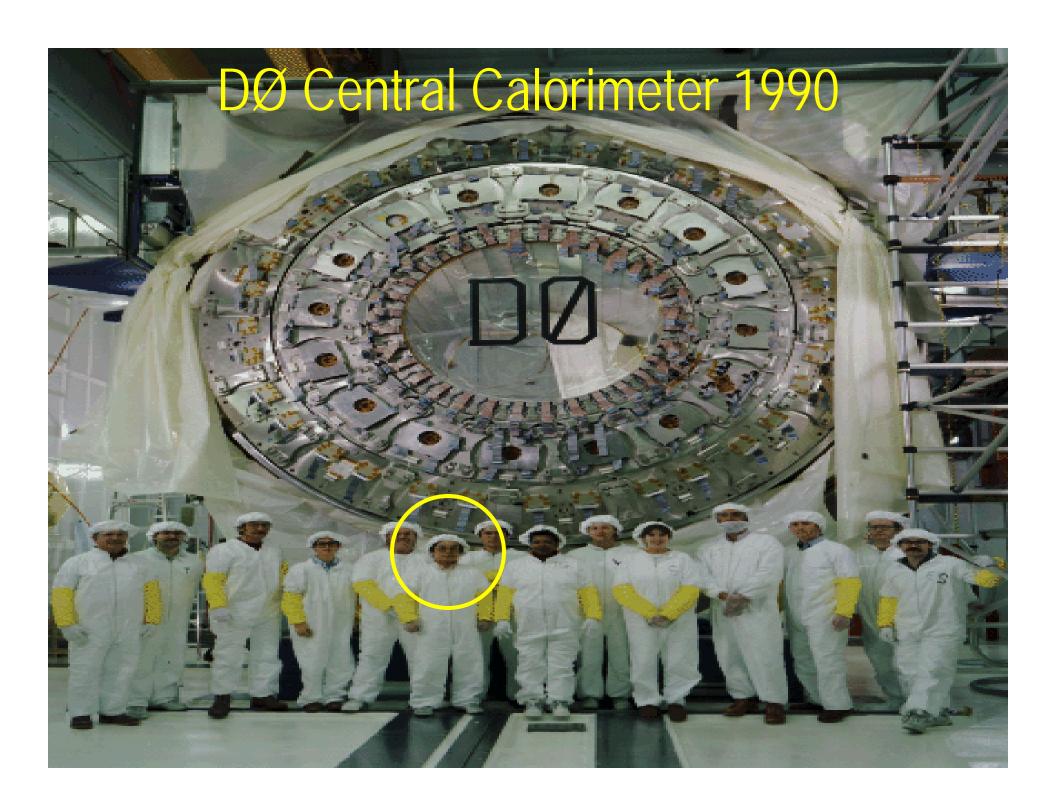
ATLAS Detector

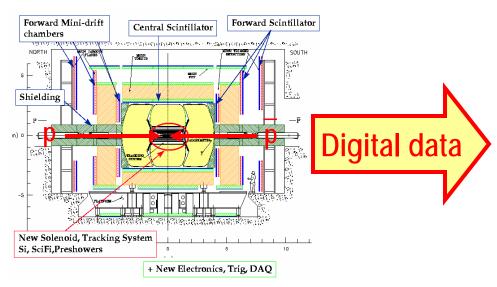


- Weighs 10000 tons and 10 story tall
- Can inspect 1,000,000,000 collisions/second
- Will record 100 200 collisions/second
- Records approximately 300,000,000 bytes/second
- Will record **1.5x10**¹⁵ (1,500,000,000,000,000) bytes each year (1.5 PetaByte).

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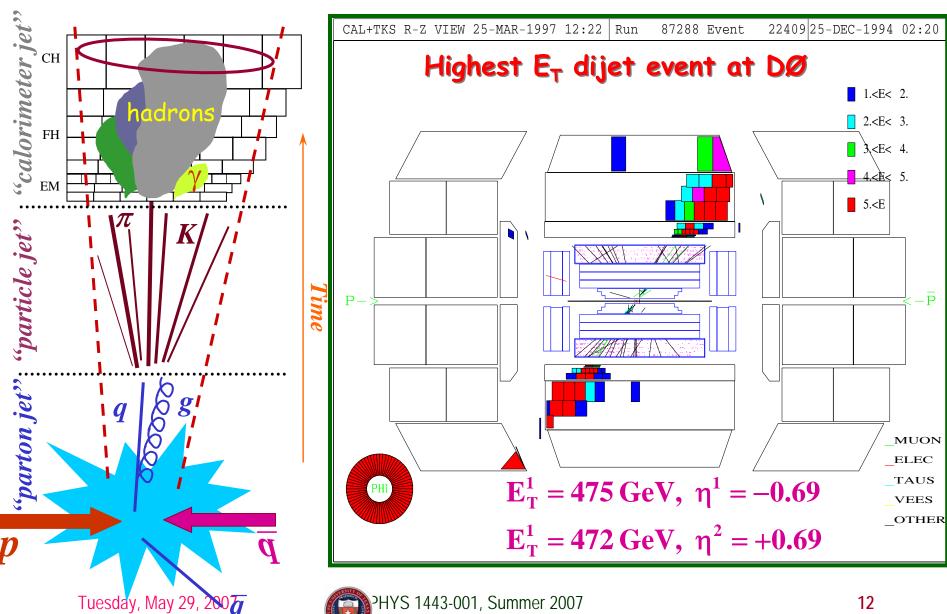






1443-001, Summer 2007 Dr. Jaehoon Yu

How does an Event Look in a Collider Detector?



Information & Communication Source

- My web page: http://www-hep.uta.edu/~yu/
 - Contact information & Class Schedule
 - Syllabus
 - Homework
 - Holidays and Exam days
 - Evaluation Policy
 - Class Style & Communication
 - Other information
- - 5 points extra credit if done by next Monday, June 4
 - 3 points extra credit if done by next Wednesday, June 6
- Office Hours: 10:00 11:00am, Mondays, Wednesdays and Thursdays or by appointments

Evaluation Policy

- Term Exams: 45%
 - Total of two exams (6/14 and 7/2)
 - Both exams will be used for the final grade
 - Each will constitute 22.5% of the total
 - Missing an exam is not permissible unless pre-approved
 - No makeup test
 - You will get an F if you miss any of the exams without a prior approval
- Lab score: 20%
- Homework: 25%
- Pop-quizzes: 10%

100%

- Extra credits: 10% of the total
 - Random attendances
 - Strong participation in the class discussions
 - Other many opportunities
- Will be on sliding scale unless everyone does very well

Homeworks

- Solving homework problems is the only way to comprehend class material
- An electronic homework system has been setup for you
 - Details are in the material distributed today and on the web
 - https://hw.utexas.edu/studentInstructions.html
 - Download homework #1 (1 problem), attempt to solve it, and submit it → You will receive a 100% credit for HW#1
 - Roster will close Friday, June 1
- Each homework carries the same weight
- ALL homework grades will be used for the final grade
- Home work will constitute <u>25% of the total</u> → A good way of keeping your grades high
- Strongly encouraged to collaborate

 Does not mean you can copy

Attendances and Class Style

Attendances:

- Will be taken randomly
- Will be used for extra credits

Class style:

- Lectures will be on electronic media
 - The lecture notes will be posted on the web <u>AFTER</u> each class
- Will be mixed with traditional methods
- Active participation through questions and discussions are
 STRONGLY encouraged → Extra credit....

Why do Physics?

Exp. To understand nature through experimental observations and measurements (Research)

Theory Establish limited number of fundamental laws, usually with mathematical expressions

Predict the nature's course

- ⇒Theory and Experiment work hand-in-hand
- ⇒Theory works generally under restricted conditions
- ⇒Discrepancies between experimental measurements and theory are good for improvements
- ⇒Improves our everyday lives, though some laws can take a while till we see amongst us

Models, Theories and Laws

- Models: A kind of analogy or mental image of a phenomena in terms of something we are familiar with
 - Often provides insights for new experiments and ideas
- Theories: More systematically improved version of models
 - Can provide quantitative predictions that are testable and more precise
- Laws: Certain concise but general statements about how nature behaves → The statement must be found experimentally valid
- Principles: Less general statements of how nature behaves
 - Has some level of arbitrariness

What do we want from this class?

- Physics is everywhere around you.
- Understand the fundamental principles that surrounds you in everyday lives...
- Identify what law of physics applies to what phenomena and use them appropriately
- Understand the impact of such physical laws
- Learn how to research and analyze what you observe.
- Learn how to express observations and measurements in mathematical languages.
- Learn how to express your research in systematic manner in writing
- I don't want you to be scared of PHYSICS!!!

Most of importantly, let us have a lot of FUN!!

Brief History of Physics

- AD 18th century:
 - Newton's Classical Mechanics: A theory of mechanics based on observations and measurements
- AD 19th Century:
 - Electricity, Magnetism, and Thermodynamics
- Late AD 19th and early 20th century (Modern Physics Era)
 - Einstein's theory of relativity: Generalized theory of space, time, and energy (mechanics)
 - Quantum Mechanics: Theory of atomic phenomena
- Physics has come very far, very fast, and is still progressing, yet we've got a long way to go
 - What is matter made of?
 - How do matters get mass?
 - How and why do matters interact with each other?
 - How is universe created?

Needs for Standards and Units

- Three basic quantities for physical measurements
 - Length, Mass, and Time
- Need a language that everyone can understand each other
 - Consistency is crucial for physical measurements
 - The same quantity measured by one must be comprehendible and reproducible by others
 - Practical matters contribute
- A system of unit called **SI** (*System Internationale*) was established in 1960
 - Length in meters (m)
 - Mass in kilo-grams (kg)
 - <u>Time</u> in seconds (s)

Definition of Base Units

SI Units	Definitions	
1 m (Length) = 100 cm	One meter is the length of the path traveled by light in vacuum during a time interval of 1/299,792,458 of a second.	
1 kg (Mass) = 1000 g	It is equal to the mass of the international prototype of the kilogram, made of platinum-iridium in International Bureau of Weights and Measure in France.	
1 s (Time)	One second is the <u>duration of 9,192,631,770</u> <u>periods of the radiation</u> corresponding to the transition between the two hyperfine levels of the ground state of the Cesium 133 (C ¹³³) atom.	

- There are prefixes that scales the units larger or smaller for convenience (see pg. 7)
- Units for other quantities, such as Kelvins for temperature, for easiness of use

Prefixes, expressions and their meanings

- deca (da): 10¹
- hecto (h): 10²
- kilo (k): 10³
- mega (M): 10⁶
- giga (G): 10⁹
- tera (T): 10¹²
- peta (P): 10¹⁵
- exa (E): 10¹⁸

- deci (d): 10⁻¹
- centi (c): 10⁻²
- milli (m): 10⁻³
- micro (μ): 10⁻⁶
- nano (n): 10⁻⁹
- pico (p): 10⁻¹²
- femto (f): 10⁻¹⁵
- atto (a): 10⁻¹⁸

International Standard Institutes

- International Bureau of Weights and Measure http://www.bipm.fr/
 - Base unit definitions:http://www.bipm.fr/enus/3_SI/base_units.html
 - Unit Conversions: http://www.bipm.fr/enus/3_SI/
- US National Institute of Standards and Technology (NIST) http://www.nist.gov/

How do we convert quantities from one unit to another?

Unit 1 = Conversion factor X Unit 2

1 inch	2.54	cm
1 inch	0.0254	m
1 inch	2.54x10 ⁻⁵	km
1 ft	30.3	cm
1 ft	0.303	М
1 ft	3.03x10 ⁻⁴	km
1 hr	60	minutes
1 hr	3600	seconds
And many	More	Here

Examples 1.3 and 1.4 for Unit Conversions

• Ex 1.3: A silicon chip has an area of 1.25in². Express this in cm².

What do we need to know?

1.25 in² = 1.25 in² ×
$$\left(\frac{2.54 \text{ cm}}{1 \text{ in}}\right)^2$$

= 1.25 in² × $\left(\frac{6.45 \text{ cm}^2}{1 \text{ in}^2}\right)$
= 1.25 × 6.45 cm² = 8.06 cm²

• Ex 1.4: Where the posted speed limit is 65 miles per hour (mi/h or mph), what is this speed (a) in meters per second (m/s) and (b) kilometers per hour (km/h)?

1 mi=
$$(5280 \text{ ft}) \left(\frac{12 \text{ in}}{1 \text{ ft}}\right) \left(\frac{2.54 \text{ cm}}{1 \text{ in}}\right) \left(\frac{1 \text{ m}}{100 \text{ cm}}\right) = 1609 \text{ m} = 1.609 \text{ km}$$

(a) 65 mi/h =
$$\left(65 \text{ mi}\right) \left(\frac{1609 \text{ m}}{1 \text{ mi}}\right) \left(\frac{1}{1 \text{ h}}\right) \left(\frac{1 \text{ h}}{3600 \text{ s}}\right) = 29.1 \text{ m/s}$$

(b) 65 mi/h =
$$(65 \text{ mi}) \left(\frac{1.609 \text{ km}}{1 \text{ mi}}\right) \left(\frac{1}{1 \text{ h}}\right) = 104 \text{ km/h}$$

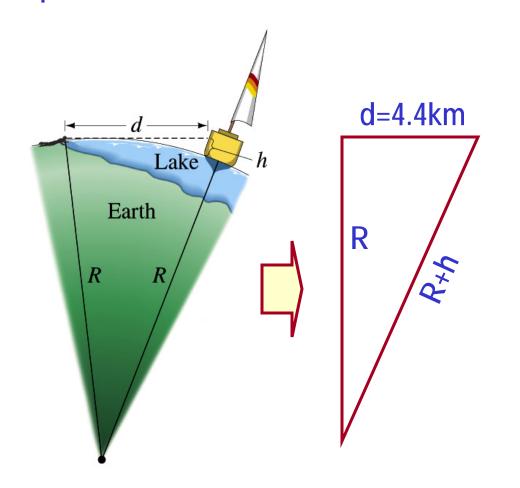
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Estimates & Order-of-Magnitude Calculations

- Estimate = Approximation
 - Useful for rough calculations to determine the necessity of higher precision
 - Usually done under certain assumptions
 - Might require modification of assumptions, if higher precision is necessary
- Order of magnitude estimate: Estimates done to the precision of 10s or exponents of 10s;
 - Three orders of magnitude: 10³=1,000
 - Round up for Order of magnitude estimate; 8x10⁷ ~ 10⁸
 - Similar terms: "Ball-park-figures", "guesstimates", etc

Example 1.8

Estimate the radius of the Earth using triangulation as shown in the picture when d=4.4km and h=1.5m.



Pythagorian theorem

$$(R+h)^2 \approx d^2 + R^2$$

 $R^2 + 2hR + h^2 \approx d^2 + R^2$
Solving for R

$$R \approx \frac{d^{2} - h^{2}}{2h}$$

$$= \frac{(4400m)^{2} - (1.5m)^{2}}{2 \times 1.5m}$$

$$= 6500km$$