

PHYS 5326 – Lecture #9 & 10

Friday, Feb. 21, 2003

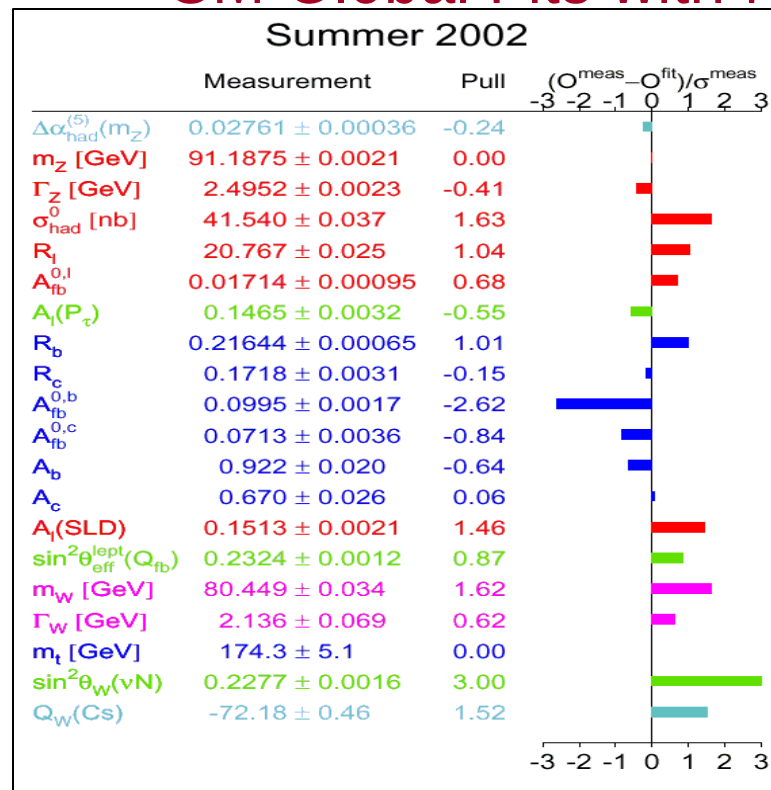
Dr. Jae Yu

1. Interpretation of $\text{Sin}^2\theta_W$ results
2. The link to Higgs
3. Neutrino Oscillation

•Next makeup class is Friday, Mar. 14, 1-2:30pm, rm 200.



SM Global Fits with New Results



LEP EWWG: <http://www.cern.ch/LEPEWWG>

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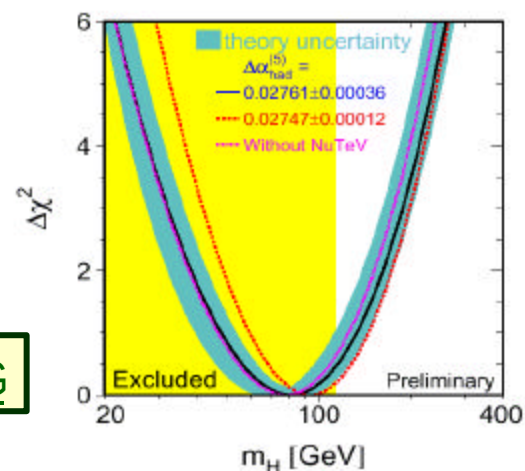
Without NuTeV

$\chi^2/\text{dof}=20.5/14$: $P=11.4\%$

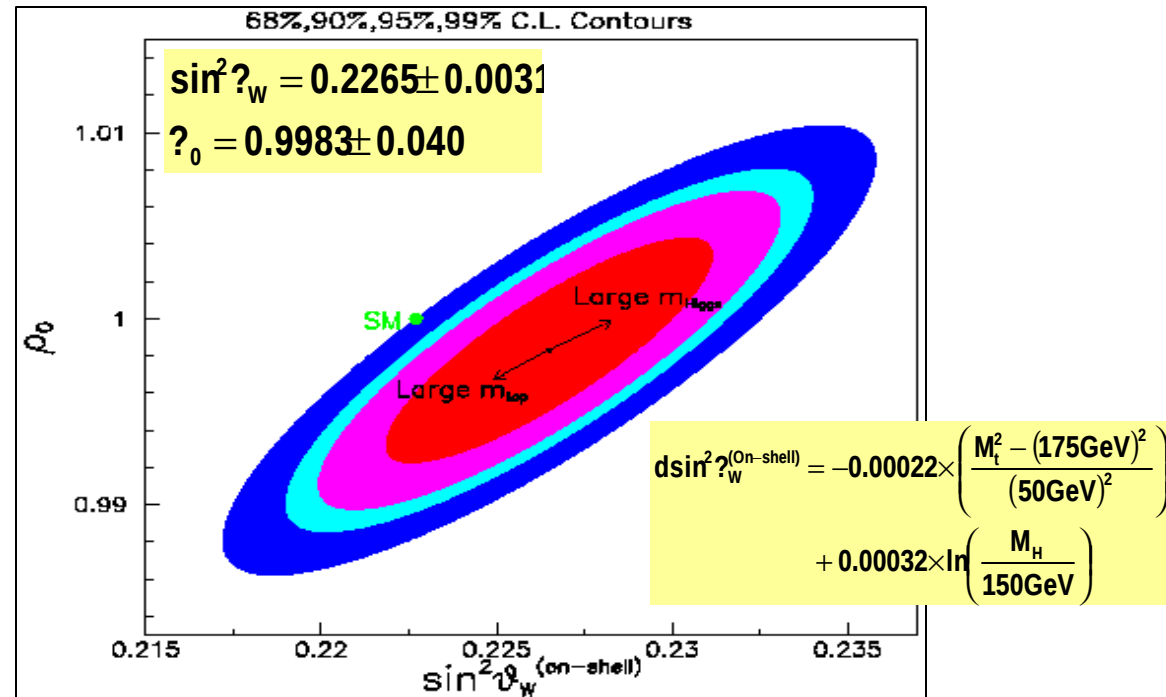
With NuTeV

$\chi^2/\text{dof}=29.7/15$: $P=1.3\%$

Confidence level in upper
 M_{higgs} limit weakens slightly.



Tree-level Parameters: ρ_0 and $\sin^2\theta_W^{(\text{on-shell})}$



- Either $\sin^2\theta_W^{(\text{on-shell})}$ or ρ_0 could agree with SM but both agreeing simultaneously is unlikely



Model Independent Analysis

- $R^{\nu}(\bar{\nu})$ can be expressed in terms of quark couplings:

$$R^{n(\bar{n})} \equiv \frac{s \left(\begin{smallmatrix} (-) \\ n \ N \rightarrow n \ X \end{smallmatrix} \right)}{s \left(\begin{smallmatrix} (-) \\ n \ N \rightarrow \ell^{-(+)} X \end{smallmatrix} \right)} = g_L^2 + r^{(-1)} g_R^2$$

Where $r \equiv \frac{s(\bar{?}N \rightarrow \ell^+ X)}{s(?N \rightarrow \ell^{-(+)} X)} \approx \frac{1}{2}$

Paschos-Wolfenstein formula can be expressed as

$$R^- = \frac{s_{NC}^? - s_{NC}^{\bar{?}}}{s_{CC}^? - s_{CC}^{\bar{?}}} = ?^2 \left(\frac{1}{2} - \sin^2 ?_w \right) = \frac{R^? - r R^{\bar{?}}}{1-r} = g_L^2 - g_R^2$$



Model Independent Analysis

- Performed a fit to quark couplings (and g_L and g_R)
 - For isoscalar target, the νN couplings are

$$g_L^2 = u_L^2 + d_L^2 = ?_0^2 \left(\frac{1}{2} - \sin^2 ?_w + \frac{5}{9} \sin^4 ?_w \right)$$

$$g_R^2 = u_R^2 + d_R^2 = ?_0^2 \frac{5}{9} \sin^4 ?_w$$

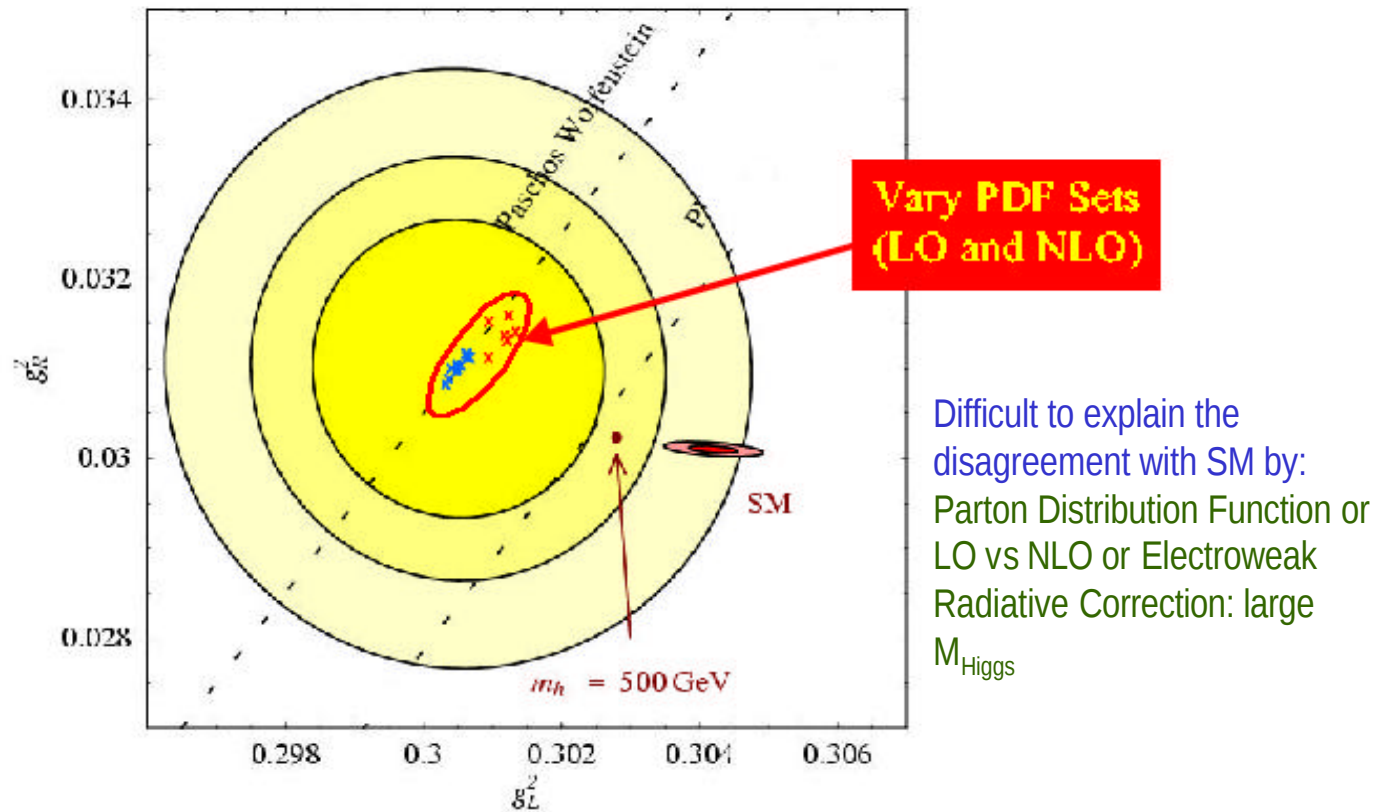
- From two parameter fit to $\mathbf{R}_n^{\text{exp}}$ and $\mathbf{R}_{\bar{n}}^{\text{exp}}$

$$g_L^2 = 0.3005 \pm 0.0014 \quad (\text{SM: } 0.3042 \leftarrow -2.6\sigma \text{ deviation})$$

$$g_R^2 = 0.0310 \pm 0.0011 \quad (\text{SM: } 0.0301 \leftarrow \text{Agreement})$$



Model Independent Analysis



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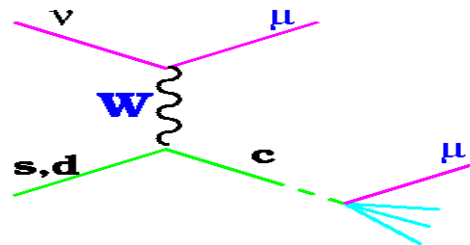
What is the discrepancy due to (Old Physics)?

- R^- technique is sensitive to q vs $\bar{q}$ differences and NLO effect
 - Difference in valence quark and anti-quark momentum fraction
- Isospin symmetry assumption might not be entirely correct
 - Expect violation about 1% $\rightarrow$ NuTeV reduces this effect by using the ratio of ν and $\bar{\nu}$ cross sections
 - $\rightarrow$ Reducing dependence by a factor of 3



What is the discrepancy due to (Old Physics)?

- s vs $\bar{s}$ quark asymmetry
 - s and $\bar{s}$ needs to be the same but the momentum could differ
 - A value of $\Delta s = x_s - x_{\bar{s}} \sim +0.002$ could shift $\sin^2\theta_W$ by -0.0026 , explaining $\frac{1}{2}$ the discrepancy (S. Davison, et. al., hep-ph/0112302)
 - NuTeV di- μ measurement shows that $\Delta s \sim -0.0027 \pm 0.0013$



Use opposite sign di- μ events to measure s and $\bar{s}$.



What is the discrepancy due to (Old Physics)?

- NLO and PDF effects
 - PDF, m_c , Higher Twist effect, etc, are small changes
- Heavy vs light target PDF effect (Kovalenko et al., hep-ph/0207158)
 - Using PDF from light target on Iron target could make up the difference → NuTeV result uses PDF extracted from CCFR (the same target)



$\nu_e \rightarrow \nu_s$ Oscillations with Large M_ν

- LSND result implicate a large Δm^2 ($\sim 10 - 100 \text{eV}^2$) solution for ν_e oscillation $\rightarrow$ MiniBooNe at FNAL is running to put the nail on the coffin
- How would this affect NuTeV $\sin^2 \theta_W$?

$$\sin^2 \theta_W = \frac{1}{2} - \frac{R^n - rR^n}{1-r} \quad \text{and} \quad R^n = \frac{N_{\text{Short}}^n - N_{n_e}^{MC}}{N_{\text{Long}}^n}$$

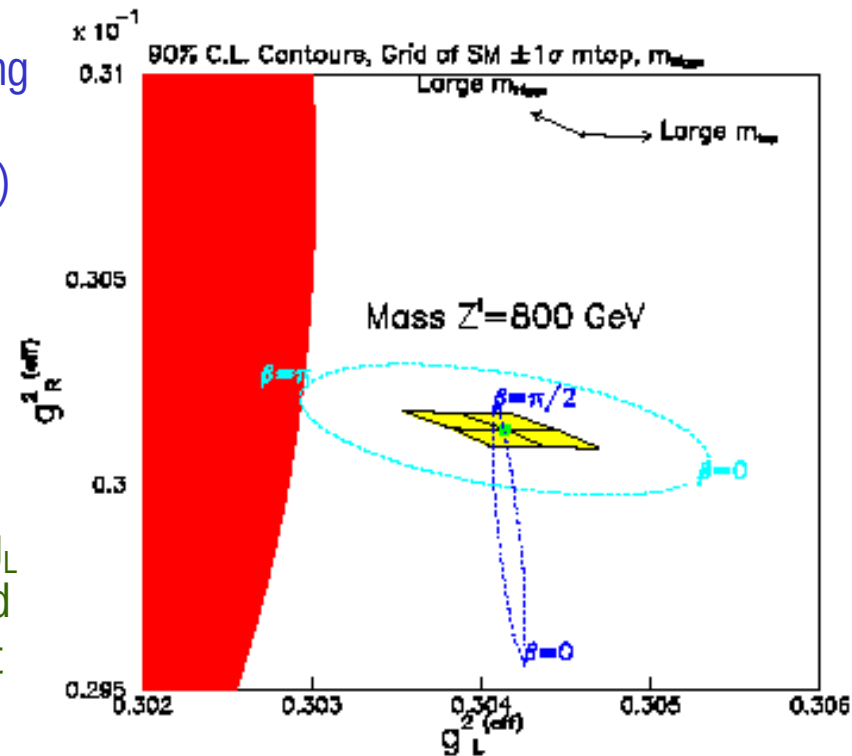
If $\nu_e \rightarrow \nu_s$ with $P_{n_e \rightarrow n_s}$ then $N_{n_e} = N_{n_e}^{MC} P_{n_e \rightarrow n_e} = N_{n_e}^{MC} (1 - P_{n_e \rightarrow n_s})$

Thus, MC will subtract more than it is in nature, causing measured R^ν to be smaller and thereby increasing $\sin^2 \theta_W$



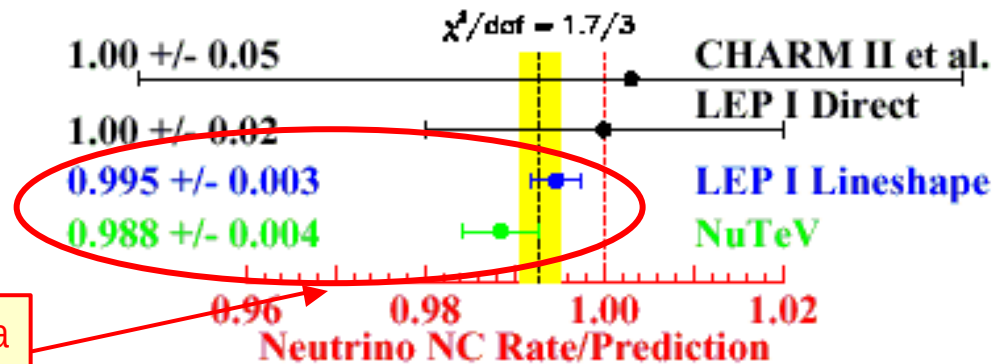
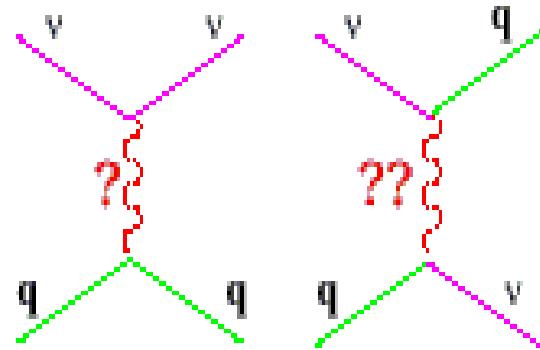
New Physics: Interactions from Extra U(1) – Z'

- Extra U(1) gauge group giving rise to interactions mediated by heavy Z' boson ($M_{Z'} \gg M_Z$)
- While couplings in these groups are arbitrary, E(6) gauge groups can provide mechanism for extra U(1) interaction via heavy Z'.
- Can give rise to g_R but not g_L which is strongly constrained by precision Z measurement



What other explanations (New Physics)?

- Heavy non-SM vector boson exchange: Z' , LQ , etc
 - Suppressed $Z\nu\nu$ (invisible) coupling
 - LL coupling enhanced than LR needed for NuTeV



Both precision data
are lower than SM

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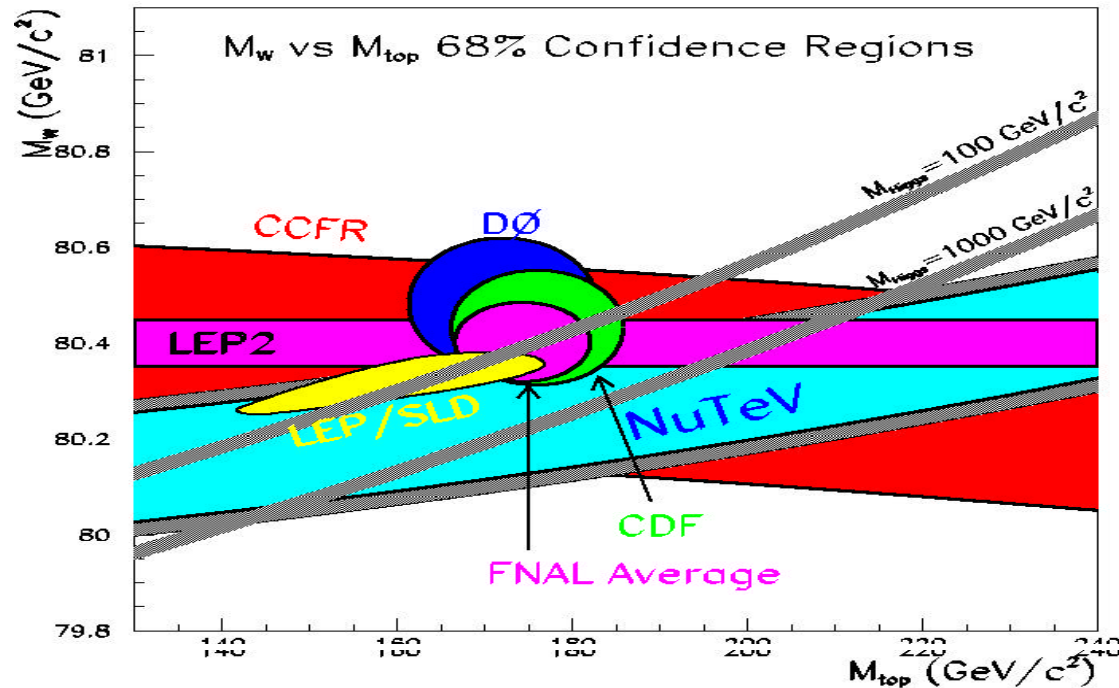
What other explanations (New Physics)?

- Propagator and coupling corrections
 - Small compared to the effect
- MSSM : Loop corrections wrong sign and small for the effect
- Many other attempts in progress but so far nothing seems to explain the NuTeV results
 - Lepto-quarks
 - Contact interactions with LL coupling (NuTeV wants $m_Z \sim 1.2\text{TeV}$, CDF/DØ: $m_Z > 700\text{GeV}$)
 - Almost sequential Z' with opposite coupling to ν

Langacker *et al.*, Rev. Mod. Phys. **64** 87; Cho *et al.*, Nucl. Phys. **B531**, 65; Zppenfeld and Cheung, hep-ph/9810277; Davidson *et al.*, hep-ph/0112302



Linking $\sin^2\theta_W$ with Higgs through M_{top} vs M_W



One-loop
correction to
 $\sin^2\theta_W$

$$d\sin^2\theta_W^{(\text{On-shell})} = -0.00022 \times \left(\frac{M_t^2 - (175\text{GeV})^2}{(50\text{GeV})^2} \right) + 0.00032 \times \ln\left(\frac{M_H}{150\text{GeV}} \right)$$

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Neutrino Oscillation

- First suggestion of neutrino mixing by B. Pontecorvo at the K^0 , K^0 -bar mixing in 1957
- Solar neutrino deficit in 1969 by Ray Davis in Homestake Mine in SD. → Called MSW effect
- Caused by the two different eigenstates for mass and weak
- Neutrinos change their flavor as they travel → Neutrino flavor mixing
- Oscillation probability depends on
 - Distance between the source and the observation point
 - Energy of the neutrinos
 - Difference in square of the masses



Neutrino Oscillation Formalism

- Two neutrino mixing case:

$$\begin{pmatrix} n_e \\ n_m \end{pmatrix} = \begin{pmatrix} \cos\theta & \sin\theta \\ -\sin\theta & \cos\theta \end{pmatrix} \begin{pmatrix} n_1 \\ n_2 \end{pmatrix} \quad \text{OR} \quad \begin{aligned} |n_e\rangle &= \cos\theta |n_1\rangle + \sin\theta |n_2\rangle \\ |n_m\rangle &= -\sin\theta |n_1\rangle + \cos\theta |n_2\rangle \end{aligned}$$

where $|n_e\rangle$ and $|n_m\rangle$ are weak eigenstates, while $|n_1\rangle$ and $|n_2\rangle$ are mass eigenstates, and θ is the mixing angle that give the extent of mass eigenstate mixture, analogous to Cabbio angle



Oscillation Probability

- Let ν_μ at time $t=0$ be the linear combination of ν_1 and ν_2 with masses m_1 and m_1 , the wave function becomes:

$$|n_m(t=0)\rangle = -\sin q |n_1\rangle + \cos q |n_2\rangle$$

- Then later time t the ν_μ wave function becomes:

$$|n_m(t)\rangle = -\sin q \exp\left[-i\left(\frac{E_1}{\hbar}\right)t\right] |n_1\rangle + \cos q \exp\left[-i\left(\frac{E_2}{\hbar}\right)t\right] |n_2\rangle$$

- For relativistic neutrinos ($E_\nu \gg m_i$), the energies of the mass eigenstates are:

$$E_k = \sqrt{p^2 + m_k^2} \cong p + \frac{m_k^2}{2p}$$



Oscillation Probability

- Substituting the energies into the wave function:

$$|n_m(t)\rangle = \exp\left[-it\left(p + \frac{m_1^2}{2E_n}\right)\right] \left[-\sin q |n_1\rangle + \cos q |n_2\rangle \exp\left[\frac{i\Delta m^2 t}{2E_n}\right] \right]$$

where $\Delta m^2 \equiv m_1^2 - m_2^2$ and $E_n \cong p$.

- Since the ν 's move at the speed of light, $t=x/c$, where x is the distance to the source of ν_μ .
- The probability for ν_μ with energy E_ν oscillates to ν_e at the distance L from the source becomes

$$P(n_\mu \rightarrow n_e) = \sin^2 2q \sin^2 \left(\frac{1.27 \Delta m^2 L}{E_n} \right)$$



Homework Assignments

- Produce an electron E_T spectrum of the highest E_T electrons in your samples
 - Due Wednesday, Feb. 26
- Complete the derivation of the probability for nm of energy E_ν to oscillate to ν_e at the distance L away from the source of ν_μ .
- Draw the oscillation probability distributions as a function of
 - Distance L for a fixed neutrino beam energy E_ν ($=5, 50, 150$ GeV)
 - E_ν for a detector at a distance L ($=1.5, 735, 2200$ km) away from the source
- Due Wednesday, Mar. 5

