

PHYS 3446 – Lecture #8

Monday, Feb. 14, 2005

Dr. Jae Yu

1. Nuclear Models

- Shell Model Predictions
- Collective Model
- Super-deformed nuclei

2. Nuclear Radiation

- Alpha Decay
- Beta Decay
- Gamma Decay



Announcements

- All of you have been given accounts at a DPCC computer
 - Please pick up your account sheet and bring it to Wednesday tutorial
- Tutorial Wednesday
 - Takes place in SH203
 - Gather in SH200 first and move to the next door
 - Your Mav-express cards will allow you access to SH203 for your projects after today
- Quiz results
 - Top score: 67
 - Average: 38.5
- First term exam
 - Date and time: 1:00 – 2:30pm, Monday, Feb. 21
 - Location: SH125
 - Covers: Appendix A (special relativity) + CH1 – CH4.4



Nuclear Models

- Liquid Droplet Model:
 - Ignore individual nucleon quantum properties
 - Assume spherical shape of nuclei
 - A core with saturated nuclear force + loosely bound surface nucleons
 - Describes BE of light nuclei reasonably well
- Fermi Gas Model:
 - Assumes nucleus as a gas of free protons and neutrons confined to the nuclear volume
 - Takes into account quantum effects w/ discrete nucleon energy levels
 - Accounts for strong spin pairing of nucleons
- Shell Model
 - Takes into account individual nucleon quantum properties
 - Needed to postulate a few potential shapes for nucleus
 - The model using spin-orbit potential seems reproduce all the desired magic numbers



Predictions of the Shell Model

- Spin-Parity of large number of odd-A nuclei predicted well
 - Nucleons obey Pauli exclusion principle → Fill up ground state energy levels in pairs
 - Ground state of all even-even nuclei have zero total angular momentum
- Single particle shell model cannot predict odd-odd nuclei spins
 - No prescription for how to combine the unpaired proton and neutron spins



Predictions of the Shell Model

- Magnetic Moment of neutron and proton are

$$\mu_p \approx 2.79 \mu_N \quad \mu_n \approx -1.91 \mu_N$$

- Intrinsic magnetic moment of unpaired nucleon to contribute to total magnetic moment of nuclei

- Deuteron

$$\mu_D = \mu_p + \mu_n = 2.79 \mu_N - 1.91 \mu_N = 0.88 \mu_N$$

- Measured value is $\mu_D = 0.86 \mu_N$

- For Boron ($^{10}\text{B}^5$), the neutrons and protons have the same level structure: $(1S_{1/2})^2(1P_{3/2})^3$, leaving one of each unpaired and one proton in angular momentum $l=1$ state $\rightarrow \mu = \frac{e\hbar}{2m_N c} l = \mu_N$

$$\mu_B = \mu_p + \mu_n + \mu_{\text{orbit}} = 2.79 \mu_N - 1.91 \mu_N + \mu_N = 1.88 \mu_N$$

- Measured value is $\mu_B = 1.80 \mu_N$

- Does not work well with heavy nuclei



Collective Model

- For heavy nuclei, shell model predictions do not agree with experimental measurements
 - Especially in magnetic dipole moments
- Measured values of quadrupole moments for closed shells differ significantly with experiments
 - Some nuclei's large quadrupole moments suggests significant nonspherical shapes
 - The assumption of rotational symmetry in shell model does not seem quite right
- These deficiencies are somewhat covered through the reconciliation of liquid drop model with Shell model
 - Bohr, Mottelson and Rainwater's collective model, 1953



Collective Model

- Assumption

- Nucleus consists of hard core of nucleons in filled shells
- Outer valence nucleons behave like the surface molecules in a liquid drop
- Non-sphericity of central core caused by the surface motion of the valence nucleon

- Thus, in collective model, the potential is a shell model with a spherically asymmetric potential

- Aspherical nuclei can produce additional energy levels upon rotation while spherical ones cannot

- Important predictions of collective model:

- Existence of rotational and vibrational energy levels in nuclei
- Accommodate decrease of spacing between first excited state and the ground level for even-even nuclei as A increases, since moment of inertia increases with A
- Spacing is largest for closed shell nuclei, since they tend to be spherical

Super-deformed Nuclei

- Nuclei tend to have relatively small intrinsic spins
- Particularly stable nuclei predicted for A between 150 and 190 with spheroidal character
 - Semi-major axis about a factor of 2 larger than semi-minor
- Heavy ion collisions in late 1980s produced super-deformed nuclei with angular momentum of $\sim 60\hbar$
- The energy level spacings of these observed through photon radiation seem to be essentially fixed
- Different nuclei seem to have identical emissions as they spin down
- Problem with collective model and understanding of strong pairing of nucleon binding energy
- Understanding nuclear structure still in progress



Nuclear Radiation: Alpha Decay

- Represents the disintegration of a parent nucleus to a daughter through an emission of a He nucleus
- Reaction equation is



- α -decay is a spontaneous fission of the parent nucleus into two daughters of highly asymmetric masses
- Assuming parent at rest, from the energy conservation

$$M_P c^2 = M_D c^2 + T_D + M_\alpha c^2 + T_\alpha$$

- Can be re-written as

$$T_D + T_\alpha = (M_P - M_D - M_\alpha) c^2 = \Delta M c^2$$



Nuclear Radiation: Alpha Decay

- Since electron masses cancel, we could use atomic mass expression

$$T_D + T_\alpha = \left(M(A, Z) - M(A-4, Z-2) - M(4, 2) \right) c^2 \equiv Q$$

- This is the definition of the disintegration energy or Q-value
 - Difference of rest masses of the initial and final states
 - Q value is equal to the sum of the final state kinetic energies
- For non-relativistic particles, KE are

$$T_D = \frac{1}{2} M_D v_D^2 \quad T_\alpha = \frac{1}{2} M_\alpha v_\alpha^2$$



Nuclear Radiation: Alpha Decay

- Since the parent is at rest, from the momentum conservation

$$M_D v_D = M_\alpha v_\alpha \quad v_D = \frac{M_\alpha}{M_D} v_\alpha$$

- If $M_D \gg M_\alpha$, $v_D \ll v_\alpha$, then $T_D \ll T_\alpha$
- We can write the relationship of KE and Q-value as

$$T_D + T_\alpha = \frac{1}{2} M_D v_D^2 + \frac{1}{2} M_\alpha v_\alpha^2 = \frac{1}{2} M_D \left(\frac{M_\alpha}{M_D} v_\alpha \right)^2 + \frac{1}{2} M_\alpha v_\alpha^2$$

$$T_D + T_\alpha = T_\alpha \frac{M_\alpha + M_D}{M_D}$$

$$T_\alpha = \frac{M_D}{M_\alpha + M_D} Q$$

- T_α is unique for the given nuclei
- Direct consequence of 2-body decay of a rest parent



Nuclear Radiation: Alpha Decay

- KE of the emitted α must be positive
- Thus for an α -decay to occur, it must be an exothermic process $\Delta M \geq 0$, $Q \geq 0$
- For massive nuclei, the daughter's KE is

$$T_D = Q - T_\alpha = \frac{M_\alpha}{M_\alpha + M_D} Q = \frac{M_\alpha}{M_D} T_\alpha \ll T_\alpha$$

- Since $M_\alpha/M_D \approx 4/(A-4)$, we obtain

$$T_\alpha \approx \frac{A-4}{4} Q \quad T_D \approx \frac{4}{A} Q$$



Nuclear Radiation: Alpha Decay

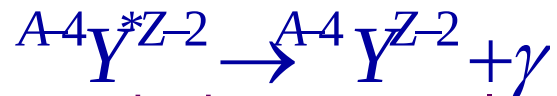
- Most energetic α -particles produced alone
 - Parent nucleus decays to the ground state of a daughter and produces an α -particle whose KE is the entire Q value

- Less energetic ones accompany photons – mostly delayed...

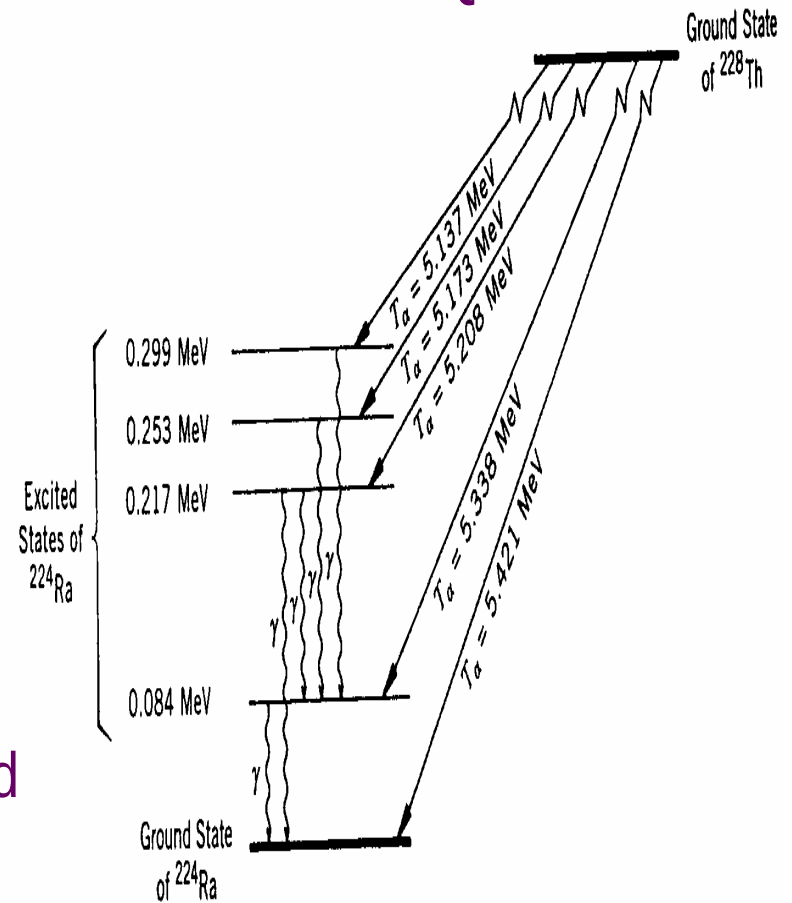
- Indicates quantum energy levels
- Parent decays to an excited state of the daughter after emitting an α



- Daughter then subsequently de-excite by emitting a photon



- Difference in the two Q values correspond to photon energy



Monday, Feb. 14, 2005



PHYS 3446, Spring 2005
Jae Yu

Nuclear Radiation: α -Decay Example

- $^{240}\text{Pu}_{94}$ decay reaction is



- α particles observed with 5.17MeV and 5.12 MeV

- Since $Q = \frac{A}{A-4} T_\alpha$

- We obtain the two Q-values

$$Q_1 \approx \frac{240}{236} 5.17 \text{MeV} = 5.26 \text{MeV} \quad Q_2 \approx \frac{240}{236} 5.12 \text{MeV} = 5.21 \text{MeV}$$

- Which yields photon energy of $E_\gamma = \Delta Q = Q_1 - Q_2 = 0.05 \text{MeV}$
- Consistent with experimental measurement, 45KeV
- Indicates the energy level spacing of order 100KeV for nuclei
 - Compares to order 1eV spacing in atomic levels



Nuclear Radiation: β -Decays

- Three kinds of β -decays

- Electron emission

- Nucleus with large N_n
 - Proton number increases by one



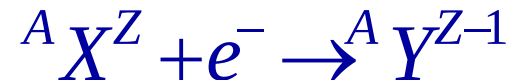
- Positron emission

- Nucleus with many protons
 - Proton number decreases by one



- Electron capture

- Nucleus with many protons
 - Absorbs a K-shell atomic electron
 - Proton number decreases by one
 - Causes cascade x-ray emission from the transition of remaining atomic electrons



- For β -decay: $\Delta A=0$ and $|\Delta Z|=1$

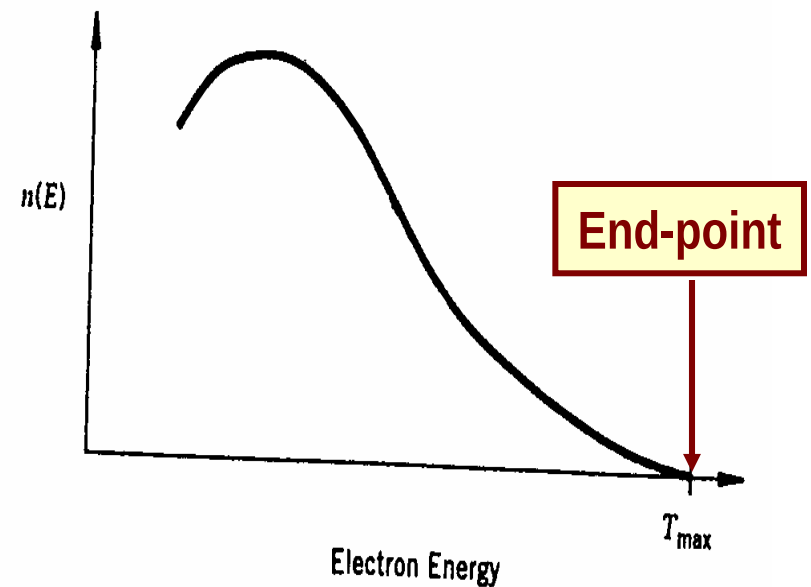


Nuclear Radiation: β -Decays

- Initially assumed to be 2-body decay
- From the conservation of energy

$$E_X = E_Y + E_{e^-} = E_Y + T_e + m_e c^2$$

- Since the lighter electron carries most the energy
 $T_e = (E_X - E_Y - m_e c^2) = (m_X - m_Y - m_e) c^2 - T_Y = Q - T_Y \approx Q$
- Will result in a unique values as in α -decay.
- In reality, electrons emitted with continuous E spectrum with an end-point given by the formula above
- Energy conservation is violated!!!!**



Monday, Feb. 14, 2005



PHYS 3446, Sr
Jae Yu

Nuclear Radiation: β -Decays

- Angular momentum is also in trouble
- In β -decays total number of nucleons is conserved
- Electrons are fermions with spin $\hbar/2$
- Independent of any changes of an integer orbital angular momentum, the total angular momentum cannot be conserved
- Angular momentum conservation is violated!!!



Nuclear Radiation: β -Decays

- Pauli proposed an additional particle emitted in β -decays
 - No one saw this particle in experiment
 - Difficult to detect
 - Charge is conserved in β -decay
 - Electrically neutral
 - Maximum energy of electrons is the Q values
 - Massless
 - Must conserve the angular momentum
 - Must be a fermion with spin $\hbar/2$
- This particle is called neutrino (by Feynman) and expressed as ν



Nuclear Radiation: Neutrinos

- Have anti-neutrinos $\bar{\nu}$, just like other particles
- Neutrinos and anti-neutrinos are distinguished by magnetic moment
 - Helicity is used to distinguish them $H \propto \vec{p} \cdot \vec{s}$
 - Left-handed (spin and momentum opposite direction) anti-electron-neutrinos are produced in β -decays
 - Right-handed electron-neutrinos are produced in positron emission
 - e^- is a particle and e^+ is an anti-particle
 - ν_e is a particle and $\bar{\nu}_e$ is an anti-particle



Assignments

1. End of the chapter problems: 3.2
2. Derive the following equations:
 - Eq. 4.8 starting from conservation of energy
 - Eq. 4.11 both the formula
- Due for these homework problems is next Wednesday, Feb. 23.

