

PHYS 3446 – Lecture #9

Wednesday, Feb. 23, 2005

Dr. Jae Yu

1. Nuclear Radiation

- Beta Decay
- Gamma Decay

2. Application of Nuclear Physics



Announcements

- Please be sure to start doing the analysis using the computer skills you have learned
 - Generate some plots, especially those Venkat asked you to since they provide good benchmark
- Term exam results
 - Average: 63.2/100 including extra credit
 - Top score: 95
 - Good job!
 - This exam constitutes 15% of the total
- There are other opportunities
 - One more term exam on Mar. 21: 15%
 - Homework: 15%
 - Lab: 15%
 - Class project: Paper (20%) + Oral presentation (10%)
 - Extra credit: up to 10% of the total



Nuclear Radiation: β -Decays

- Three kinds of β -decays

- Electron emission

- Nucleus with large N_n
 - Proton number increases by one



- Positron emission

- Nucleus with many protons
 - Proton number decreases by one



- Electron capture

- Nucleus with many protons
 - Absorbs a K-shell atomic electron
 - Proton number decreases by one
 - Causes cascade x-ray emission from the transition of remaining atomic electrons



- For β -decay: $\Delta A=0$ and $|\Delta Z|=1$

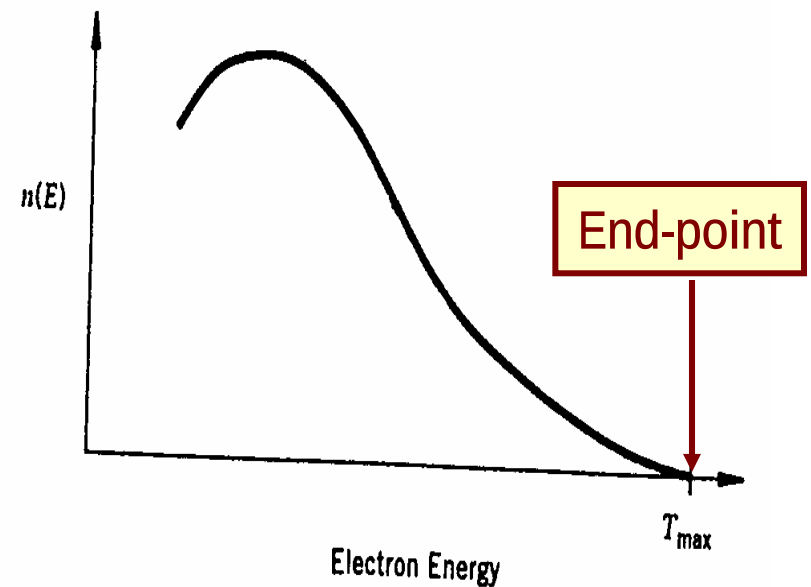


Nuclear Radiation: β -Decays

- Initially assumed to be 2-body decay
- From the conservation of energy

$$E_X = E_Y + E_{e^-} = E_Y + T_e + m_e c^2$$

- Since the lighter electron carries most the energy
 $T_e = (E_X - E_Y - m_e c^2) = (m_X - m_Y - m_e) c^2 - T_Y = Q - T_Y \approx Q$
- Will result in a unique values as in α -decay.
- In reality, electrons emitted with continuous E spectrum with an end-point given by the formula above
- Energy conservation is violated!!!!**



Nuclear Radiation: β -Decays

- Angular momentum is also in trouble
- In β -decays total number of nucleons is conserved
- Electrons are fermions with spin $\hbar/2$
- Independent of any changes of an integer orbital angular momentum, the total angular momentum cannot be conserved
- Angular momentum conservation is violated!!!



Nuclear Radiation: β -Decays

- Pauli proposed an additional particle emitted in β -decays
 - No one saw this particle in experiment
 - Difficult to detect
 - Charge is conserved in β -decays
 - Electrically neutral
 - Maximum energy of electrons is the Q values
 - Massless
 - Must conserve the angular momentum
 - Must be a fermion with spin $\hbar/2$
- This particle is called neutrino (by Feynman) and expressed as ν



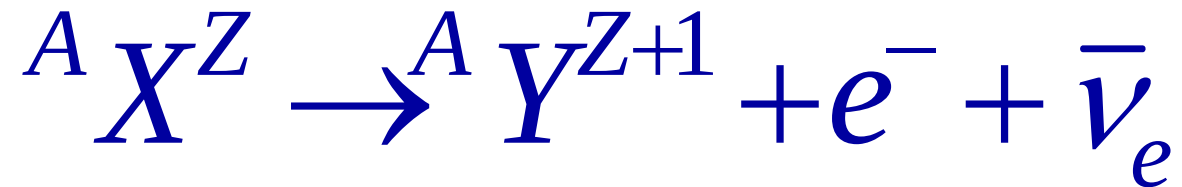
Nuclear Radiation: Neutrinos

- Have anti-neutrinos $\bar{\nu}$, just like other particles
- Neutrinos and anti-neutrinos are distinguished through the spin projection on momentum
 - Helicity is used to distinguish them $H \propto \vec{p} \cdot \vec{s}$
 - Left-handed (spin and momentum opposite direction) anti-electron-neutrinos are produced in β -decays
 - Right-handed electron-neutrinos are produced in positron emission
 - e^- is a particle and e^+ is an anti-particle
 - ν_e is a particle and $\bar{\nu}_e$ is an anti-particle

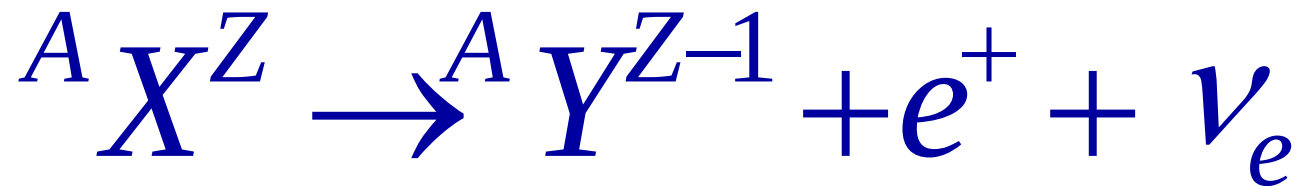


β -Decays with neutrinos

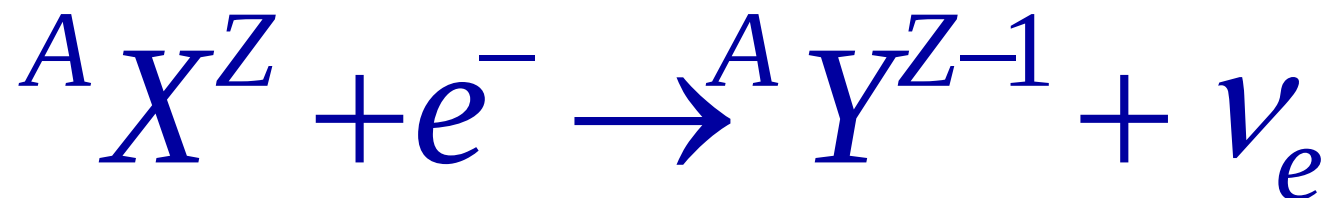
- Electron emission



- Positron emission



- Electron capture



β -Decays with neutrinos

- If the parent nucleus decays from rest, from the conservation of energy

$$M_p c^2 = T_D + M_D c^2 + T_{e^-} + m_e c^2 + T_{\bar{\nu}_e} + m_{\bar{\nu}_e} c^2$$

- Thus the Q-value of a β -decay can be written

$$T_D + T_{e^-} + T_{\bar{\nu}_e} = (M_p - M_D - m_e - m_{\bar{\nu}_e}) c^2 = \Delta M c^2 = Q$$

- Electron emission can only occur if $Q > 0$
- Neglecting all small atomic BE, e emission can occur if

$$Q = (M(A, Z) - M(A, Z+1) - m_{\bar{\nu}_e}) c^2$$
$$\approx (M(A, Z) - M(A, Z+1)) c^2 \geq 0$$



β -Decays with neutrinos

- Daughter nucleus is much heavier than e or ν , the small recoil energy of daughter can be ignored.
- Thus we can obtain $T_{e^-} + T_{\bar{\nu}_e} \approx Q$
- This means that the energy of electron is not unique and can be any value in the range $0 \leq T_{e^-} \leq Q$
- The maximum electron kinetic energy can be Q .
- The same can apply to the other two β -decays



Particle Numbers

- Baryon numbers: A quantum number assigned to baryons (particles consists of quarks)
 - Mostly conserved in many interactions
 - Baryons: +1
 - Anti-baryons: -1
 - Protons and neutrons are baryons with baryon number +1 each
- Lepton numbers: A quantum number assigned to leptons (electrons, muons, taus and their corresponding neutrinos)
 - Leptons: +1
 - Anti-leptons: -1
 - Must be conserved at all times under SM in each species



Lepton Numbers

- Three charged leptons exist in nature with their own associated neutrinos

$$\begin{pmatrix} e^- \\ \nu_e \end{pmatrix} \quad \begin{pmatrix} \mu^- \\ \nu_\mu \end{pmatrix} \quad \begin{pmatrix} \tau^- \\ \nu_\tau \end{pmatrix}$$

- These three types of neutrinos are distinct from each other
 - muon neutrinos never produce other leptons than muons or anti-muons

$$\nu_\mu + {}^A X^Z \rightarrow {}^A Y^{Z+1} + \mu^-$$

$$\nu_\mu + {}^A X^Z \not\rightarrow {}^A Y^{Z+1} + e^-$$

$$\nu_\mu + {}^A X^Z \not\rightarrow {}^A Y^{Z+1} + \tau^-$$



Lepton Numbers

For electron neutrinos

$$\nu_e + {}^A X^Z \rightarrow {}^A Y^{Z+1} + e^-$$

$$\nu_e + {}^A X^Z \not\rightarrow {}^A Y^{Z+1} + \mu^-$$

$$\nu_e + {}^A X^Z \not\rightarrow {}^A Y^{Z+1} + \tau^-$$

For tau neutrinos

$$\nu_\tau + {}^A X^Z \rightarrow {}^A Y^{Z+1} + \tau^-$$

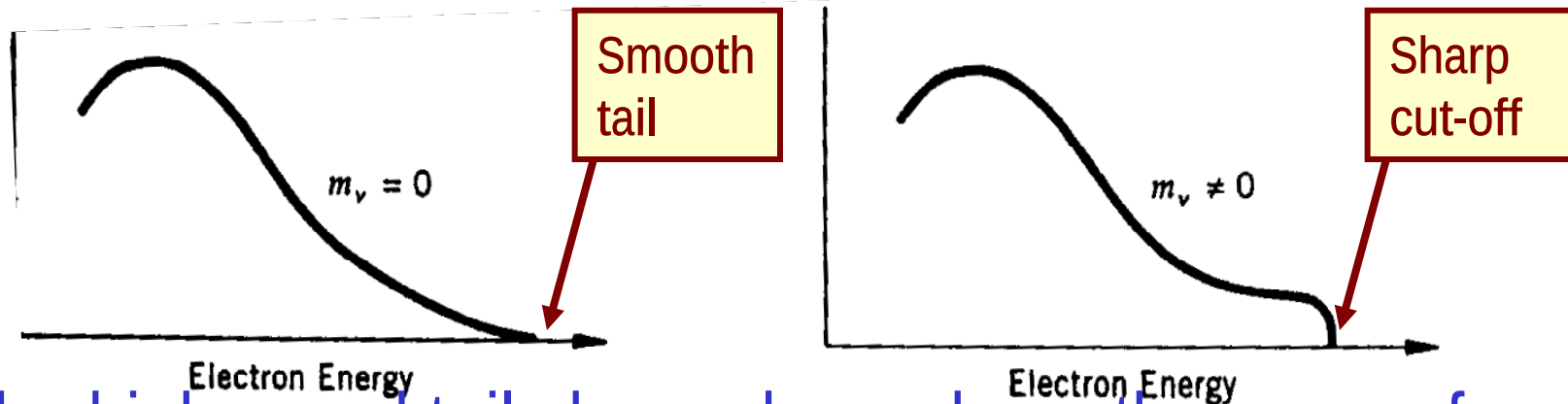
$$\nu_\tau + {}^A X^Z \not\rightarrow {}^A Y^{Z+1} + e^-$$

$$\nu_\tau + {}^A X^Z \not\rightarrow {}^A Y^{Z+1} + \mu^-$$



Neutrino Mass

- What does neutrino mass do to the β -spectrum?



- The higher end tail shape depends on the mass of neutrinos
 - β -spectrum could be used to measure the mass of neutrinos
 - Very sensitive to the resolution on the device
 - Most stringent limit on neutrino mass is $m_\nu < 2\text{eV}/c^2$
- Non-zero mass of neutrino means
 - Neutrino Oscillation: Mixing of neutrino species

Weak Interactions

- β -decay can be written in a nucleon level as:

$$n \rightarrow p + e^- + \bar{\nu}_e \quad p \rightarrow n + e^+ + \nu_e \quad p + e^- \rightarrow n + \nu_e$$

- Since neutrons are heavier than protons, they can decay to a proton in a free space
 - On the other hand, protons are lighter than neutrons therefore they can only undergo a β -decay within a nucleus
 - Life time of a neutron is about 900sec
 - This life time is a lot longer than nuclear reaction time scale 10^{-23} s or EM scale 10^{-16} s.
- This means that a β -decay is a nuclear phenomenon that does not involve strong nuclear or EM forces
- Fermi postulated a new weak force responsible for β -decay



Weak Interactions

- Weak forces are short ranged
 - Occurs in nuclear domain
 - Weakness of the strength is responsible for long life time
- Nucleus does not contain electrons
 - Electrons in β -decays must come from somewhere else
 - The electron must come at the time of decay just like the photons from the transition of atomic electrons
 - β -decay can be considered to be induced by a weak force
- The transition probability per unit time, the width, can be calculated from perturbation theory using Fermi's Golden rule

$$P = \frac{2\pi}{\hbar} |H_{fi}|^2 \rho(E_f)$$

- Where the weak interaction Hamiltonian is

$$H_{fi} = \langle f | H_{wk} | i \rangle = \int d^3x \psi_f^*(x) H_{wk} \psi_i(x)$$



Weak Interactions

- Based on β -decay reaction equations, the H_{wk} must be a four fermionic states
 - H_{wk} proposed by Fermi is a four-fermion interaction or current-current interaction
 - Relativistic
 - Agreed rather well with experiments for low energy β -decays
- Parity violation
 - There are only left-handed neutrinos and right-handed anti-neutrinos
 - A system is parity invariant if it does not change under reflection of spatial coordinates
 - The spin $\vec{r} \rightarrow -\vec{r}$, $\vec{p} \rightarrow -\vec{p} \Rightarrow \vec{L} = \vec{r} \times \vec{p} = (-\vec{r}) \times (-\vec{p}) = \vec{L}$
 - The handedness, helicity, changes upon the spatial reflection since direction of motion changes while the spin does not
 - Since there is no right handed neutrinos, parity must be violated in weak interactions



Gamma Decays

- When a heavy nuclei undergo alpha and beta decays, the daughters get into an excited state
 - Must either break apart
 - Or emit another particle
 - To bring the daughter into its ground state
- Typical energies of photons in γ -decays are a few MeV's
 - These decays are EM interactions thus the life time is on the order of 10^{-16} sec.
- Photons carry one unit of angular momentum
 - Parity is conserved in this decay



Assignments

1. End of the chapter problems: 4.4 and 4.5
2. Reading assignment: Chapter 5
3. Due for these assignments is Wednesday, Mar. 2

